BIG BANG TO BIOSIGNATURES: THE LUVOIR MISSION CONCEPT Aki Roberge (Study Scientist) Brad Peterson & Debra Fischer (STDT chairs) Julie Crooke (Study Manager) Matt Bolcar (Chief Engineer) Shawn Domagal-Goldman (Deputy Study Scientist) Decadal Studies Pause & Learn #2 June 1, 2017

Crab Nebula with HST ACS/WFC What is LUVOIR? Credit: NASA / ESA Large UV / Optical / Infrared Surveyor (LUVOIR) A space telescope concept in tradition of Hubble Broad science capabilities Far-UV to Near-IR bandpass ~ 8 – 16 m aperture diameter Suite of imagers and spectrographs Serviceable and upgradable Hubble-like guest observer program "Space Observatory for the 21st Century" Ability to answer questions we have not yet conceived

SCIENCE PROGRESS

LUVOIR online science simulation tools <u>http://asd.gsfc.nasa.gov/luvoir/tools/</u>

Large UV/Optical/Infrared Surveyor (LUVOIR)							
	National Goddard Sp	Aeronautics and Space Administration	Astrophysics Science Division - Sciences and Exploration				
LUVOIR Large UV/Optical/Infrared Surveyor							
	Home Science	On-Line Simulation Tools					
	LUVOIR Flyer	This page links to performance simulation and visualization tools for the LUVOIR mission, a future					
	Technology Seminars	ultraviolet / optical / near-infrared observatory concept. These widgets are experimental. If they are not working, email Jason Tumlinson (STScI). For the Planetary Spectrum Generator, email Geronimo Villanueva (GSFC).					
	Events						
	Meet the Team	HDI Photometric ETC	Coronagraphic Spectra of Exoplanets				
	Working Groups	Basic exposure time calculator for optical photometry in multi-band images.	Simulate optical/near-IR reflection spectra of various exoplanets with realistic noise.				
	Documents	LUMOS Spectroscopic ETC	Multiplanet Vield Tool				
	Images & Videos	Simple exposure time calculator for UV spectroscopy.	Tool for visualizing yields of observed exoplanets				
	Simulation Tools		(of various types) as function of basic mission parameters.				
	Contacts	UV MOS Visualizer	Planetary Spectrum Generator				
	For Science For Press Twitter	See the impact of UV multi-object spectroscopy on the study of stellar clusters and their feedback.	Advanced tool for simulating spectra of Solar System bodies (with LUVOIR and other telescopes).				
	Facebook	High-Resolution Imaging Examples of astronomical objects viewed with different sized telescopes.					

Imaging performance simulations



Galaxy at z = 2 with HST

Galaxy at z = 2 with 12-m LUVOIR

Credit: G. Snyder / J. Tumlinson

LUVOIR Study Update for Pause and Learn #2

Imaging Earth 2.0



Reality check



Exoplanets imaged as function of aperture

LUVOIR: The Large UltraViolet Optical Infrared Surveyor

Multiplanet Yields



Kopparapu, Stark, et al., in prep.

Solar System imaging





Pluto with HST

Pluto with 15-m LUVOIR

Credit: W. Harris



Roth et al. (2014)

UV hydrogen emission

Credit: G. Ballester (LPL)

Late-breaking ...

Jupiter from JUNO at ~ 15 km/pixel resolution Comparable to LUVOIR 15-m (~ 20 km)

Science progress summary

Science goals and needed observations defined

Instrument performance characteristics described **Table of Contents**

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1st draft of Interim Report outline complete

Writing has begun

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Interim report schedule (in reverse order)

Action	Date / Duration	Comments	
Deliver interim report to HQ	Early Dec 2017		
Revise whole report and finalize	Month of Nov	Make it pretty	
Reviews	Month of Oct	Senior advisors, Aerospace Corp, GSFC red team	
Deliver complete "reviewer's draft"	Oct 1, 2017		
STDT review of whole report	Mid Sept	At STDT meeting #6	
Finalize complete "reviewer's draft"	Month of Sept	Make it complete and uniform	
Individual writing assignments due	Mid Aug		
Assess report progress	July 31 – Aug 2	At STDT meeting #5 (joint w/ HabEx)	
Writing !	3.5 months		
Finalize report outline	Late April	Start at STDT meeting #4	

TECHNICAL PROGRESS

LUVOIR architectures overview

Architecture A

Integrated Design Center runs Jan 2017 - July 2017

- 15-m telescope
 - First Instrument Design Lab (IDL) run complete Jan 2017
- Four instruments
 - 1. Coronagraph A
 - 2. LUMOS A
 - 3. High-Definition Imager
 - 4. POLLUX

LUVOIR architectures overview

Architecture B

Integrated Design Center runs Sept 2017 - July 2018

- ~ 9-m telescope
 - First Instrument Design Lab (IDL) run in Sept 2017
- Three instruments
 - 1. Coronagraph B
 - 2. LUMOS B
 - 3. ONIRS

Observational challenge

Faint planets next to bright stars

Solution

Optical / Near-IR Coronagraph

Contrast < 10⁻¹⁰ to observe exoEarths Multi-resolution spectroscopy Bandpass: 0.2 µm to 2.5 µm Tech development via WFIRST coronagraph

First IDL run complete (March 2017)





Observational challenge

No UV through Earth's atmosphere

Solution

LUMOS

Multi-object spectroscopy (R = 500 - 45,000)

Bandpass: 100 nm to 400 nm

UV imaging

Major upgrade of HST STIS

First IDL run complete (May 2017)





HST STIS UV instrument

Observational challenge

Imaging wide fields at high resolution

Solution

High-Definition Imager

2 x 3 arcmin field-of-view
Bandpass: 0.2 μm to 2.5 μm
High precision astrometry capability (measure planet masses, etc.)
Major upgrade of HST WFC3

First IDL run complete (Feb 2017)





Observational challenge

Measuring warm molecules present in Earth's atmosphere

<u>Solution</u>

Optical / Near-IR Spectrograph

Multiple resolutions up to R ~ 10⁵ High photometric precision (transits) Possibly high precision RV capability Ground-based analogs in develop.

First IDL run in Spring 2018



Credit: Natasha Batalha



ESPRESSO spectrograph for VLT (Credit: ESO)



POLLUX: a European contribution to the LUVOIR mission study

- POLLUX is a concept for a UV spectro-polarimeter with high resolution point-source capability ($R \sim 10^5$)
- Complementary to the LUMOS instrument
- To be defined & designed by a consortium of 10 European institutions, with leadership/support from CNES
- The conceptual study conducted by CNES could serve as a support for a future ESA contribution

Preliminary Architecture A visualization



LUVOIR Architecture A (15-m)

Credit: A. Jones (GSFC)

Technology gap list (team version)

Technology	TRL	Criticality
Ultra-stable Opto-mechanical Systems	2	Enabling
Closed-loop Segment Phasing	4	
Vibration Isolation	4	
Wavefront Sensing & Control	3	
Thermal Sensing & Control	4	
Mirror Segments	5	
High-contrast Segmented Aperture Coronagraphs	4	Enabling
Coronagraph Architecture	4	
Deformable Mirrors	4	
Post-processing	3	
Ultra-Low Noise Large Format NIR Detectors:	3	Enabling
Ultra-Low Noise Large Format VIS Detectors:	4	Enabling
Ultra-Low Noise Large Format UV Detectors:	4	Enhancing
Mirror Coatings	3	Enabling

BROADER INVOLVEMENT: COMMUNITY, OTHER STUDIES, INDUSTRY, INTERNATIONAL

STDT voting members





Debra Fischer Brad Peterson (Ohio State / STScI) (Yale)





Jacob Bean (Chicago)

Lee Feinberg

(NASA GSFC)



Daniela Calzetti (U Mass)

Kevin France

(Colorado)



Rebekah Dawson (Penn State)

Olivier Guyon

(Arizona)



Courtney Dressing (Caltech)



Walter Harris (Arizona / LPL)



Mark Marley (NASA Ames)



David Redding (JPL)

June 1, 2017

Leonidas Moustakas (JPL)



Jane Rigby (NASA GSFC)



John O'Meara (St. Michael's)

Aki Roberge

(NASA GSFC)



Vikki Meadows (Washington)



David Schiminovich (Columbia)



Ilaria Pascucci (Arizona)



Britney Schmidt (Georgia Tech)



Marc Postman

(STScl)

Karl Stapelfeldt (JPL)



Jason Tumlinson (STScl)

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LUVOIR Study Update for Pause and Learn #2



Laurent Pueyo (STScl)

Face-to-face meetings

5th meeting July 31 – Aug 2, 2017 at Caltech

Observers welcome at all LUVOIR meetings & weekly telecons

Large UV/Optical/IR Surveyor (LUVOIR)

Science and Technology Definition Team Study Office, and friends

> LUVOIR STDT Meeting #1 Goddard Space Flight Center, Greenbelt MD May 9 - 10, 2016

Community involvement

Six Working Groups

- Exoplanets (21 members)
- Cosmic Origins (26 members)
- Solar System (10 members)
- Simulations (9 members)
- Communications (8 members)
- Technology (~ 65 members)

Five Instrument Teams

• Include both STDT and community participants



Community involvement

Activities at Jan 2017 AAS meeting

- Splinter meeting for hands-on work with simulation tools
- Tools demos at STScI booth in poster hall
- Three talks at NASA Hyperwall
- STDT presence at NASA booth

Splinter session at HST/JWST 5 conference in Venice (Mar 2017)

Numerous talks at conferences and universities

Community involvement LUVOIR webinars (~ monthly) Public website: <u>http://asd.gsfc.nasa.gov/luvoir/</u>



LUVOIR Study Update for Pause and Learn #2

Coordination with other studies

Three STDT members shared with HabEx STDT

Monthly LUVOIR / HabEx leadership telecons

Joint STDT meetings with HabEx team

- First joint meeting Nov 10, 2016
- Second joint meeting Aug 2, 2017

Some standard slides developed jointly with HabEx team (one to follow)

LUVOIR engineers helping HabEx with microshutters & integral field spectrographs (supported by GSFC)

Difference between LUVOIR and HabEx ?

Both LUVOIR and HabEx have two primary science goals

- Habitable exoplanets & biosignatures
- Broad range of general astrophysics
- The two architectures will be driven by difference in focus
 - For LUVOIR, both goals are on equal footing. LUVOIR will be a general purpose "great observatory", a successor to HST and JWST in the ~ 8 - 16 m class
 - HabEx will be optimized for exoplanet imaging, but also enable a range of general astrophysics. It is a more focused mission in the ~ 4 - 8 m class

Similar exoplanet goals, differing in quantitative levels of ambition

- HabEx will *explore* the nearest stars to "search for" signs of habitability & biosignatures via direct detection of reflected light
- LUVOIR will *survey* more stars to "constrain the frequency" of habitability & biosignatures and produce a statistically meaningful sample of exoEarths

The two studies will provide a continuum of options for a range of futures





Coordination with other studies

LUVOIR representation on ExEP Standards Team

Telecon w/ LUVOIR, HabEx, & OST leadership (Feb 16, 2017)

Info on development of Cooperative Agreement Notices for industry involvement shared with OST & Lynx teams Industry involvement

LUVOIR Cooperative Agreement Notice released (Dec 2016)

Two responses selected for 1-year, cost-shared studies (Mar 2017)

- Northrop Grumman, Ball Aerospace, & Harris Corp. consortium
- Lockheed Martin

Kick-off coordination meetings with both industry teams (Apr 2017)

International involvement

8 non-voting reps. of international space agencies on LUVOIR STDT

 ESA, France, Canada, UK, Spain, Japan, Denmark, Germany

POLLUX European instrument study

- Participants from 10 European institutions
- Study funded and led by French Space Agency (CNES)

Kavli IAU Workshop July 2017 in Leiden

 "Global Coordination of Ground and Space Astrophysics: Future Space-Based Optical/UV/IR Telescopes"

LESSONS LEARNED & MOVING FORWARD

What's going well

STDT engagement

Effectiveness of face-to-face STDT meetings

Design progress

Communication with HQ program scientists

• Bi-weekly study office / HQ telecons

HQ reportage appropriate and not burdensome

Challenges

STDT travel logistics

- NRESS contractor unresponsive (getting worse)
- Becoming critical issue for STDT members
- How are the other teams handling their meeting logistics? Is there a better contractor?
- Tough to design complex instruments in compressed IDL studies (1 week runs)
 - Will rely on LUVOIR Engineering Team and Instrument Teams to adjust designs, fill gaps, perform analysis

Suggestions / requests (1 of 2)

One-size-fits-all approach to reports not favored

Request flexibility in how each team designs their report(s)

Request for Decadal Studies table(s) in NASA booth area at Jan 2018 AAS

 Instead of having to stock / staff both COR and ExEP tables (like last time)

Suggestions / requests (2 of 2)

Solicit more team input on Program Office technology assessment

 So POs are up to date on current design approaches, which have been rapidly changing

What numbers to assume for launch vehicles with 5-m fairings?

• We'd like HQ to provide this, like they did for SLS