

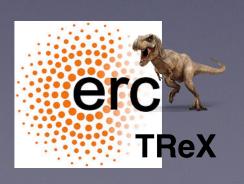
The GRB-GW Past, Present and Future

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Neil Gehrels Memorial Meeting Washington DC, May 20-21 2018



Subject: Lorentz factor of GRB 060218

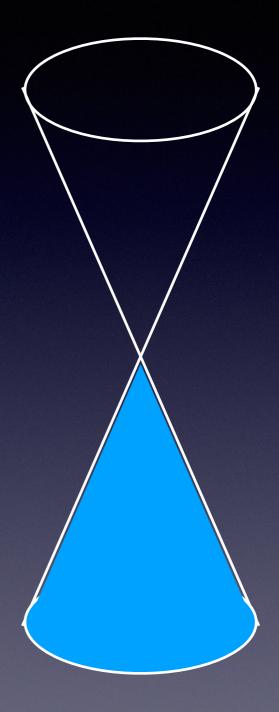
Dear Tsvi,

I noticed that in the initial astro-ph versions of the attached paper [on GRB060218] an initial Lorentz factor of ~15 was given in the abstract. In the current version, I don't see a definite number anywhere in the paper (maybe I missed it). What changed in your thinking on the Lorentz factor for this burst?

Best Regards, Neil



The Past



- 1974 r-process Nucleosynthesis from DNS mergers: *Lattimer & Schramm*
- 1975 PSR1913+16: Hules & Taylor
- 1975 GW from DNS: (Smarr and Blandford)
- 1982 1990 Shift in GW search from focus on SNe to focus on mergers (*Thorne*)
- 1989 GW+GRB+r-process: Eichler et al.
- 1993 sGRB vs LGRBs: Kovelioutou et al.
- 1997 Radioactive remnant (minisupernova later called macronova and then kilonova): *Li & Paczynski*
- 2005 GRB 090509B First evidence for sGRB-BNS association. *Gehrels et al.*, *Bloom et al.*
- 2011 Radio flare following DNS mergers:
 Nakar and Piran
- 2013 -130603B first macronova candidate: Tanvir et al., Berger et al.

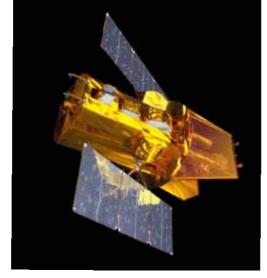


למה אנו מצפים עכשיו?

- ארוע "קרוב" (פחות ממיליארד שנות אור מאיתנו) של מיזוג כוכבי נויטרונים
 - מארוע כזה נראה
 - *גלי כבידה
 - הבזק קרינת גאמה*
 - "מקרונובה ו"יצור זהב*
 - אות רדיו מאוחר יותר *







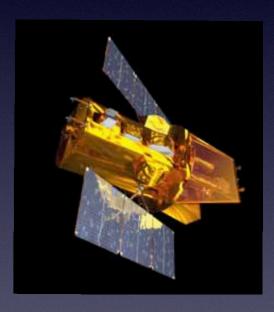


What are we expecting now?

- A nearby (< 1 Gly) binary neutron star mergers
- From such an event we will see in addition to the gravitational waves:
 - A short gamma-ray burst
 - A macronova and nucleosynthesis of "gold"
 - A radio flare

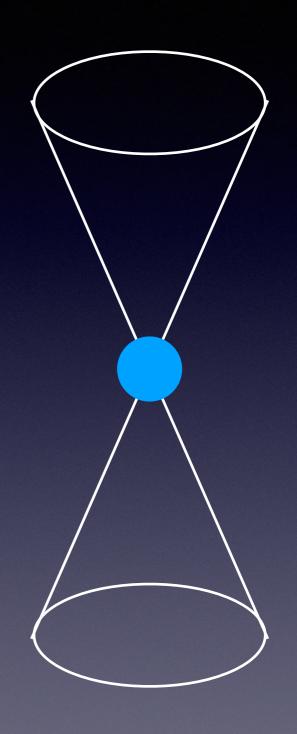








The Present
GW 170817 +
GRB170817A



GW 170817 and its EM counterparts

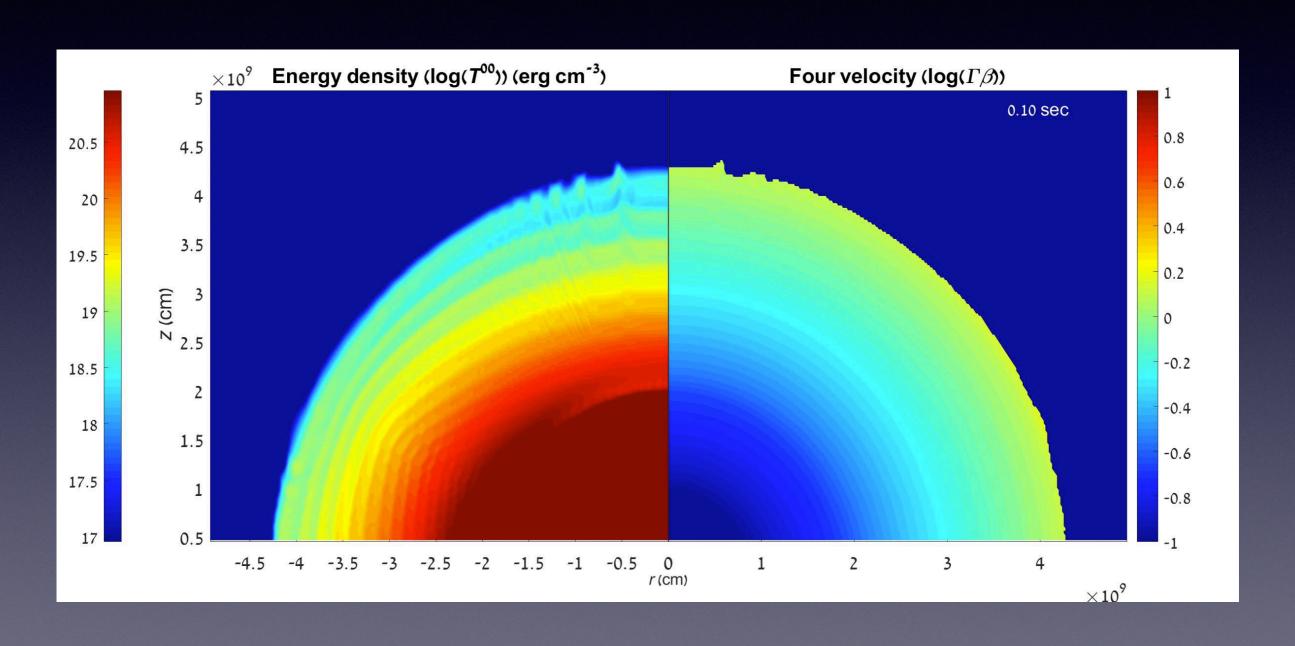
- A double NS merger + a sGRB
- An unusual low-luminosity GRB
- A macronova/kilonova => r-process nucleosynthesis
- X-ray and Radio at a relatively late time (9,16 days) inconsistent with off-axis emission of a regular sGRB
- X-ray and Radio rising as t^{0.8} over 150 days than decay.
- Is there a single picture that explains all the observations?
- The key is in the γ , the confirmation from the radio

A Low Luminosity sGRB

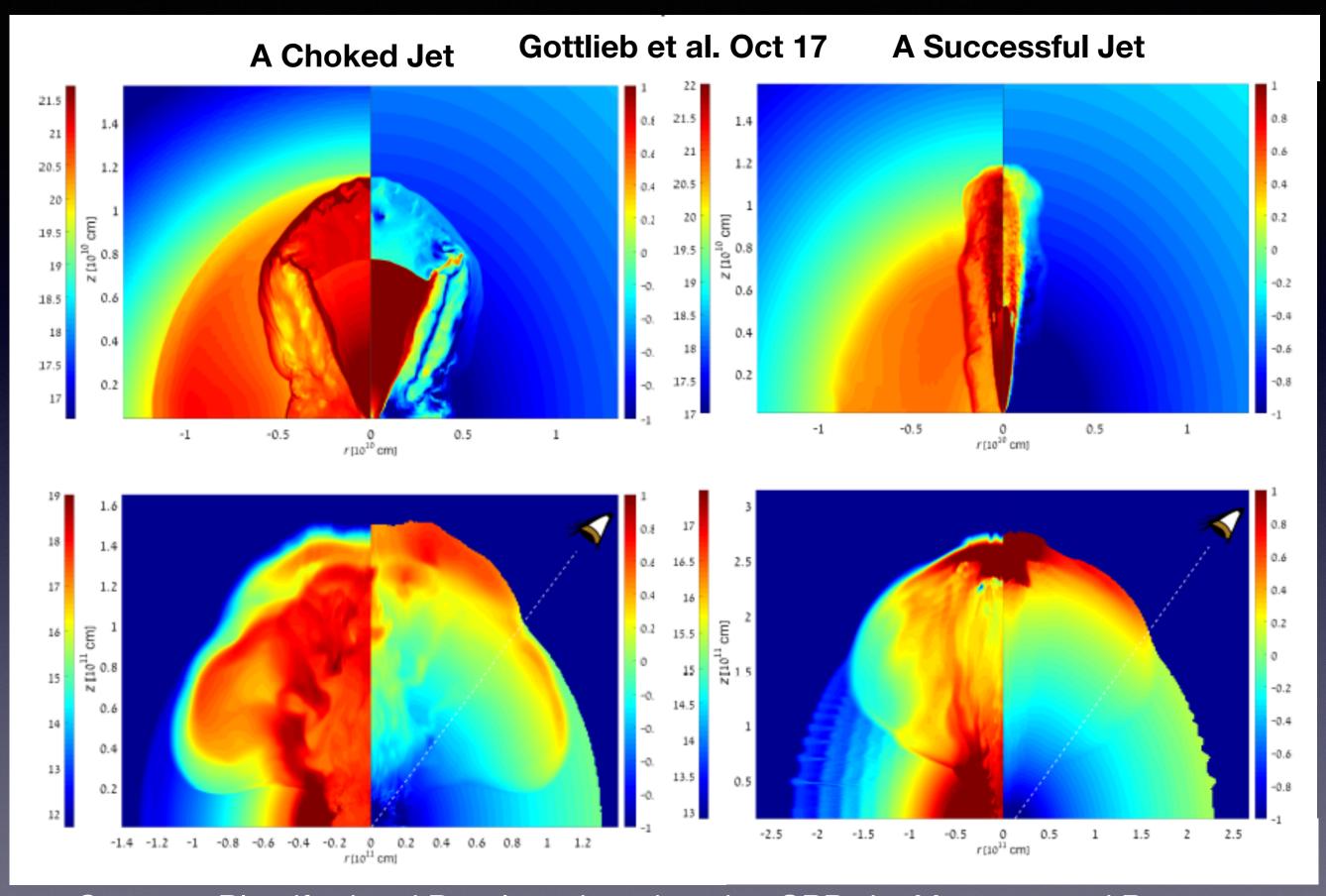
- Regular short GRBs:
 - E_{iso}~10⁵⁰-10⁵² ergs
 - Variable
 - Hard $(E_{peak} > 400 \text{ keV})$

- GRB 170817A
 - E_{iso}~10⁴⁶ ergs
 - Smooth
 - Soft (E_{peak} 180 keV)

The cocoon and the shock breakout



Credit: Ore Gottlieb

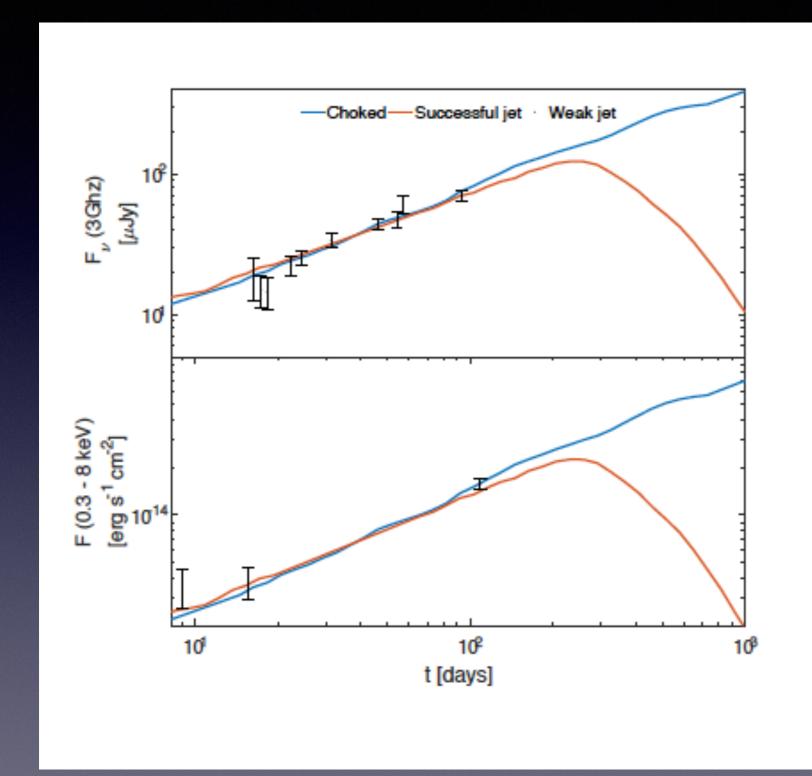


Cocoon - Blandford and Begelman introduced to GRBs by Meszaros and Rees See also: Pozanenko 17; Lazzati et a., 17 for GRB 170817A

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	Regular	Low luminosity
Long Collapsar	Non-thermal	Shock Breakout
Short Merger	Non-thermal	Shock Breakout

The radio & X-rays

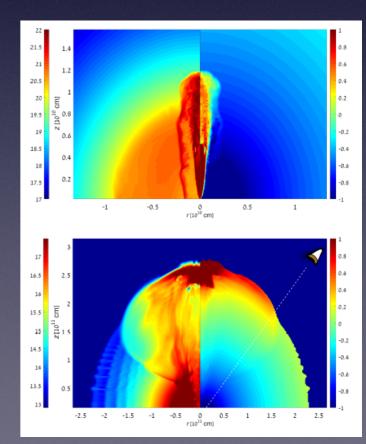


See also:
Margutti et al., 18;
Ruan et al, 18;
A'Avanzo et al., 18;
Lamb et al., 18;
Troja et al., 18;
Granot et al., 18 and others for related modeling of the radio and the x-rays.

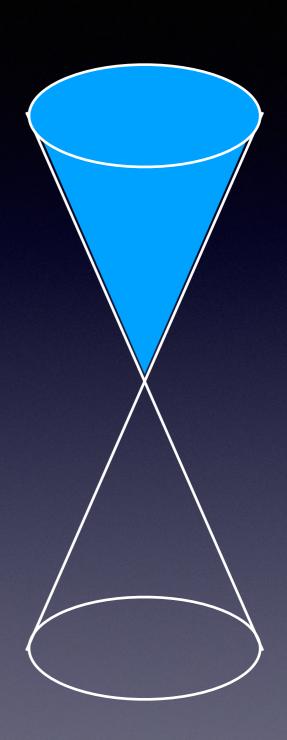
This radio signal that is produced from the same cocoons that produced the gamma-rays (From Mooley et al., 18)

A few words about words - confusion about terminology

- A <u>Structured jet</u> an outflow with E(r,θ).
 By itself this is not a physical model. A physically motivated model for a "structured jet" is a cocoon.
- A Cocoon*: The structure that arises from a propagation of a relativistic jet within external matter (also called "cloud" by some).
- ★ Coined by Blandford and Begeleman. Intoroduced to GRBs by Meszaros and Rees.



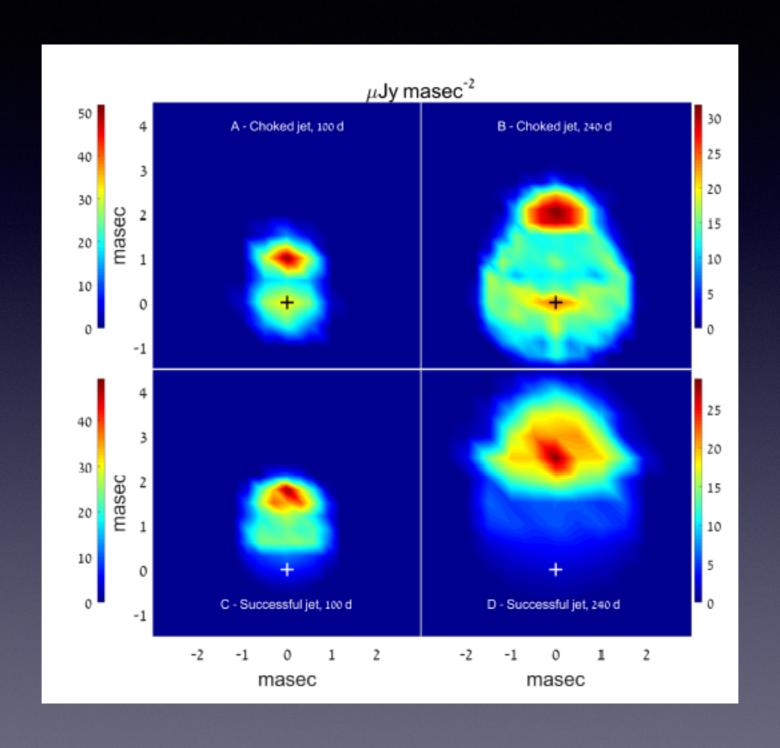




What are the questions?

- · The question is <u>not</u> "structured jet" vs. "cocoon" but:
- 1. Is the structure, $E(r,\theta)$, mostly angular or mostly radial (both arise in a "cocoon" scenario)?
- 2. What caused the structure (a "cocoon" is one option any other?).
- 3. Did the jet emerge and produce a sGRB in another direction observed by some aliens?

Was there a regular sGRB?



VLBI predictions (Gottlieb + 18)

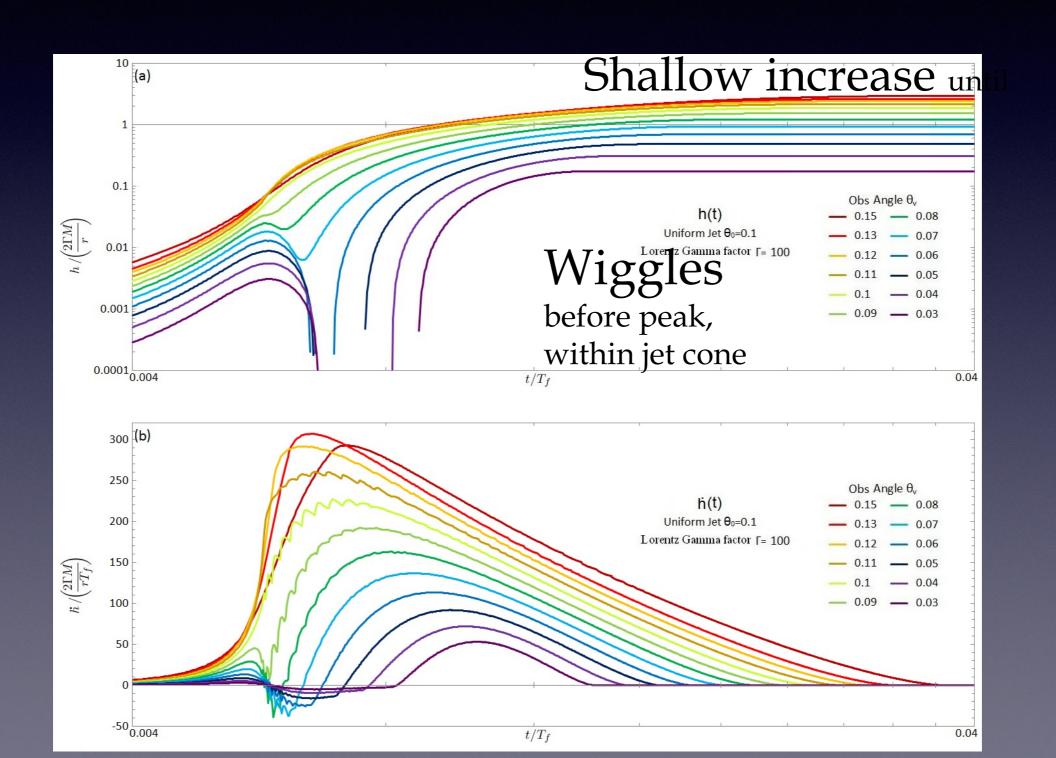
The GW - sGRB Connection

- Numerous events, including "on axis" regular sGRBs
- Most will look very different from GW 170817 (viewing angle!). The Macronova component will be most similar.
- → Statistics, nature of the remnant...

But there is more to come. With GW we will be able to observe the jet evolution.

GW from the jet

(Birnholtz & TP, 14; Leidershnier & TP 18)





Summary

- GW 170817 and sGRB 170817A confirmed the association of GW with sGRB 29 years after its prediction.
- The macronova/kilonova signal confirmed r-process nucleosynthesis in mergers 43 years after its prediction.
- A single model (Kasliwal et al., 17; Gotlieb et al. 17,...)
 explains the observed EM emission from the weak/soft
 γ-rays to the unique radio and x-rays.
- Future observations with Adv LIGO could detect GW from the accelerating (and decelerating) jet and reveal the inner working of sGRB engines!