

Chasing unidentified EGRET sources with Neil



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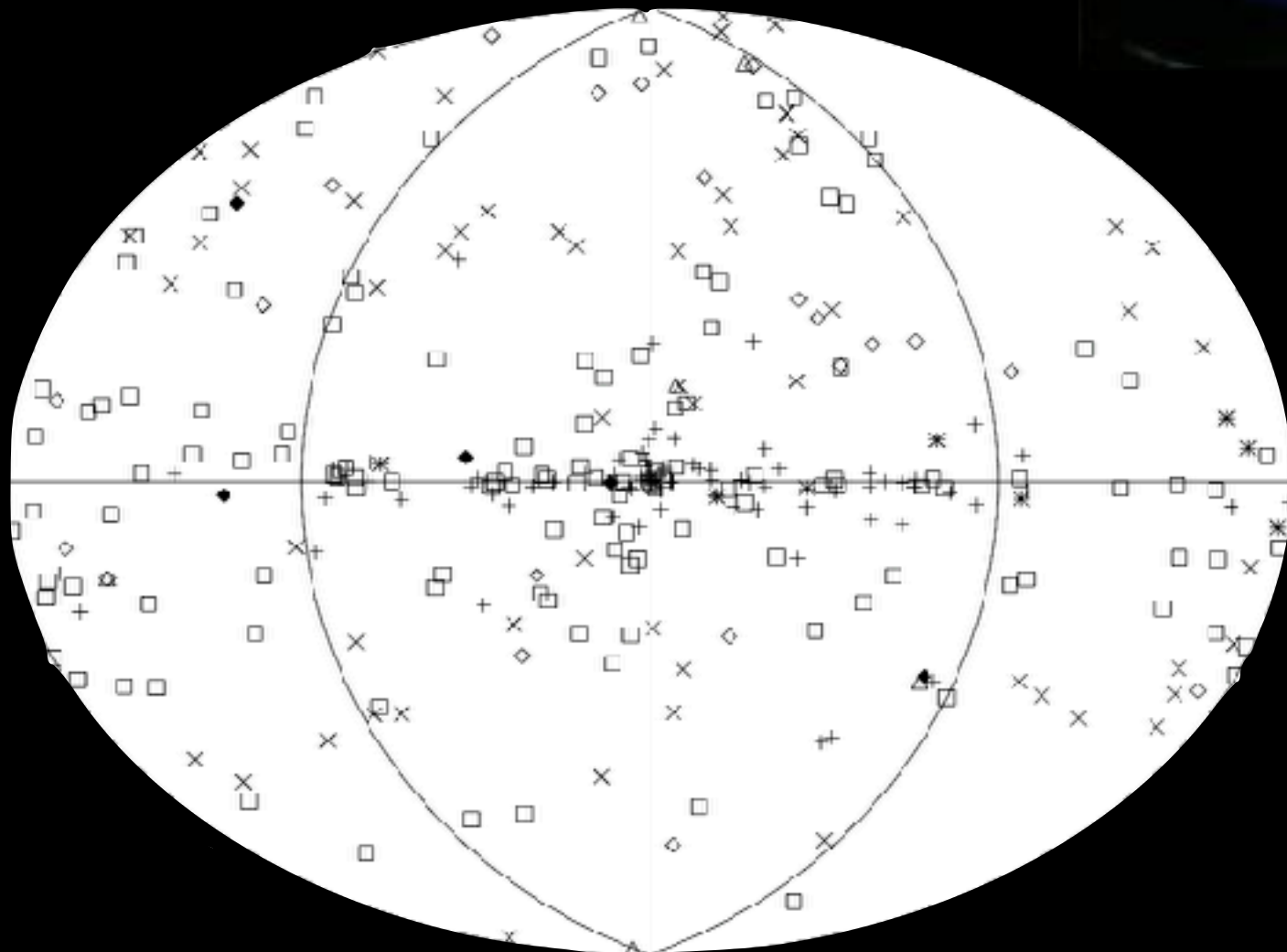
γ-ray source population studies

- Compton Gamma-Ray Observatory

- 1991-2000 (project scientist)
- CGRO-EGRET
- AGN populations

- Macomb & Gehrels 1999:

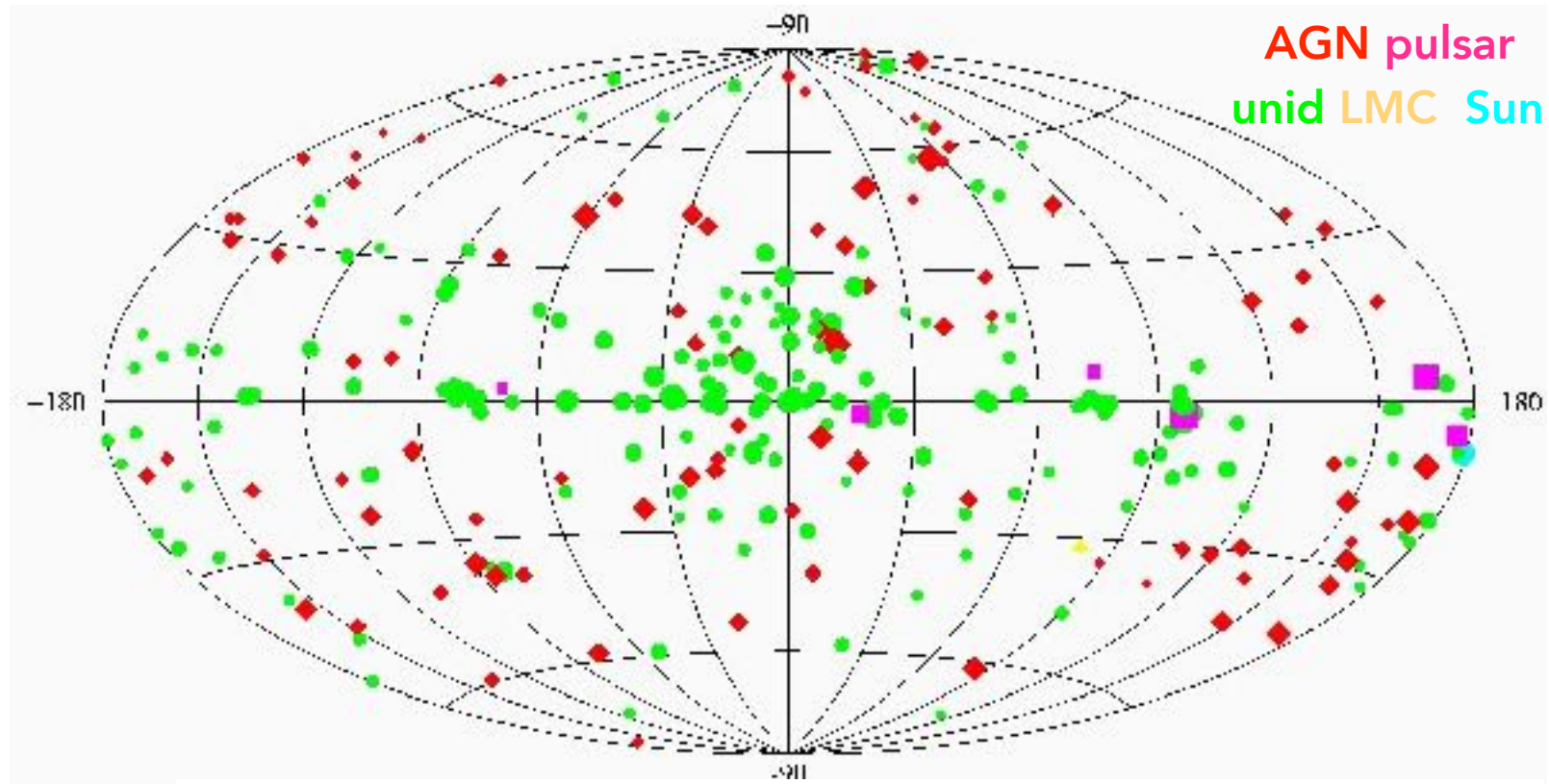
- catalogue of 309 sources
- from 50 KeV to 1 TeV




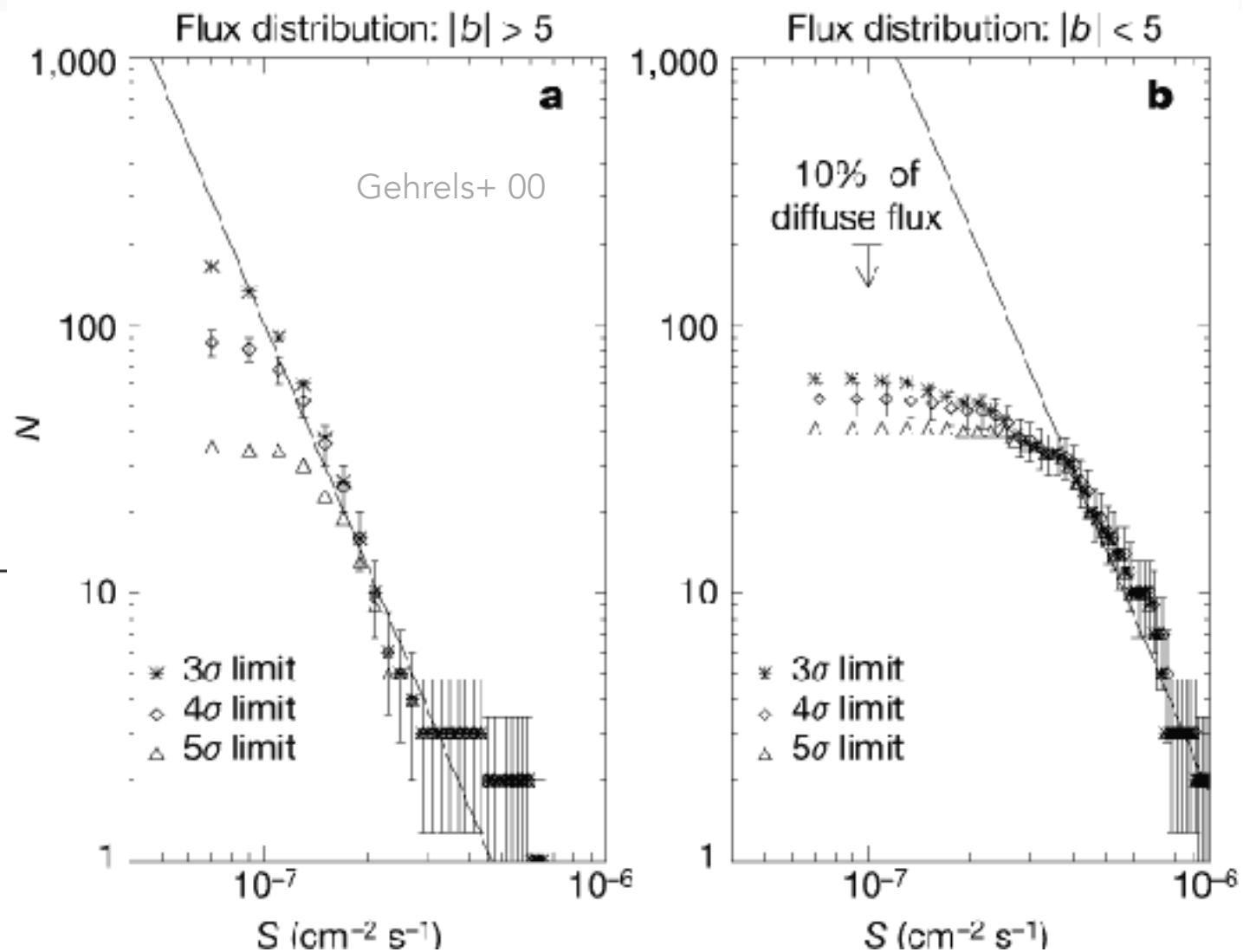
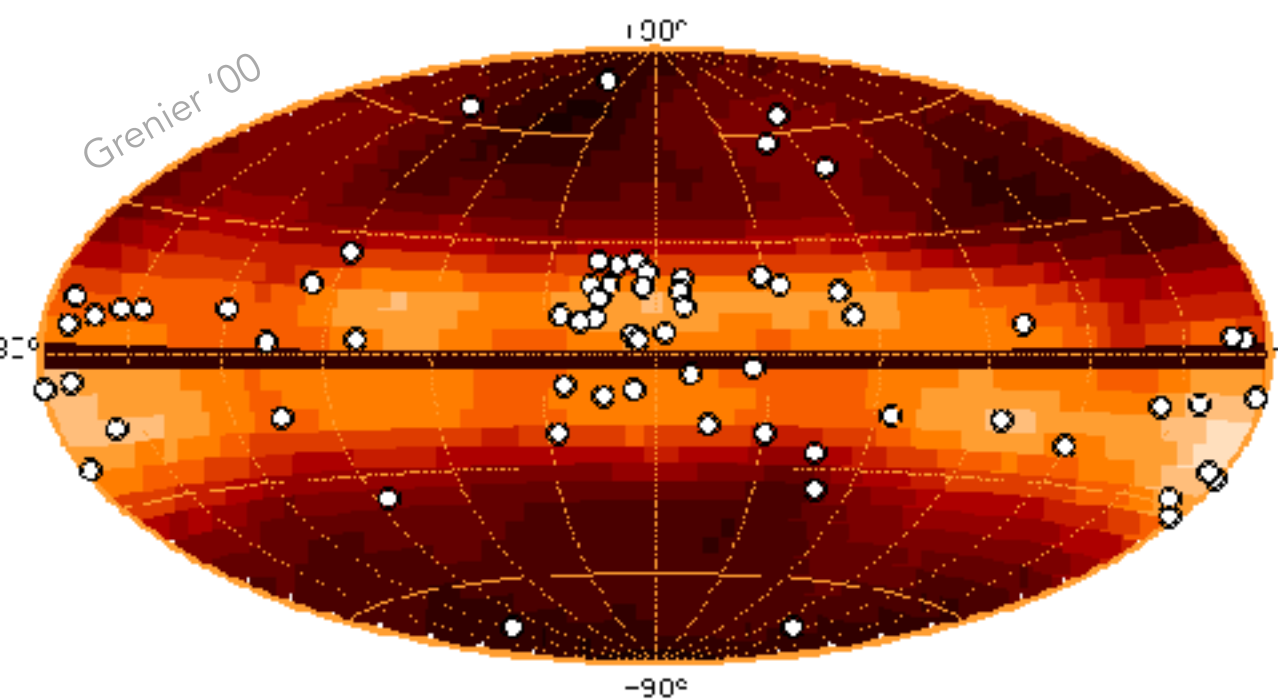
- + accreting source
- * pulsar
- ◇ AGN Seyfert
- X AGN quasar
- △ galaxy/cluster/diffuse
- unidentified
- ◆ SNR

EGRET unidentified sources

- 3rd EGRET catalog, > 100 MeV:
 - 170 unidentified sources
- Grenier 1997+99+00:
 - 45 ± 6 persistent sources associated with the Gould Belt
- Gehrels et al. 2000:
 - 20 mid-latitude sources
 - fainter and softer ($E^{-2.49 \pm 0.04}$) than at low latitudes ($E^{-2.18 \pm 0.04}$)



iso + Gould Belt model, $\alpha = -2.20$
 0.002  0.255 source/bin



origins?

● $L_{\text{iso}, > 100 \text{ MeV}} \sim 0.2-8 \cdot 10^{26} \text{ W (D/300 pc)}^2$

**3EG unidentified
young O-B3 star**

● Belt supernova relics ?

- $75 \text{ to } 95 \text{ Myr}^{-1} \text{ kpc}^{-2}$ Grenier '00
= 3 to 5 times local Gal. rate

- SN remnants,
but too extended

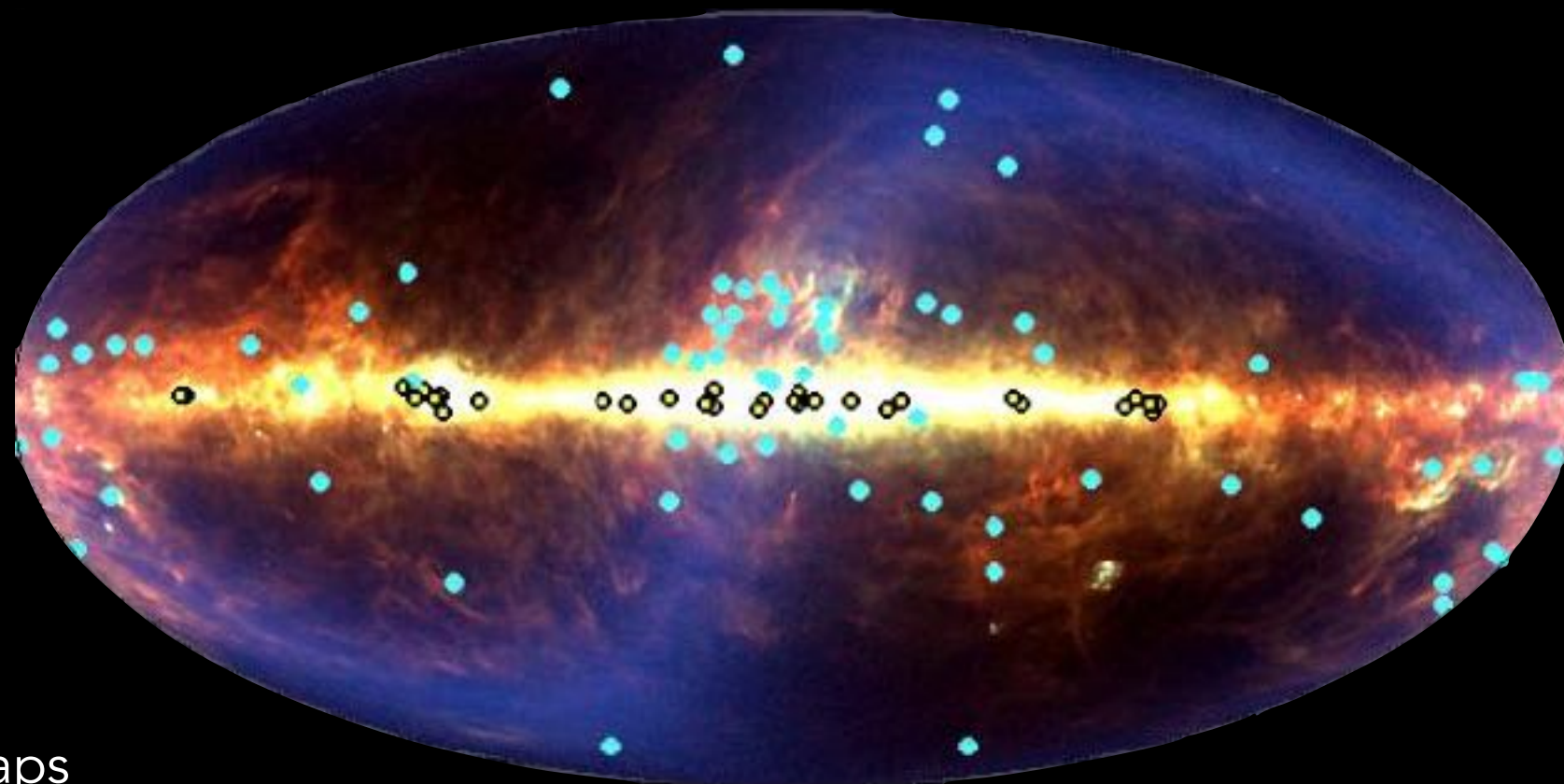
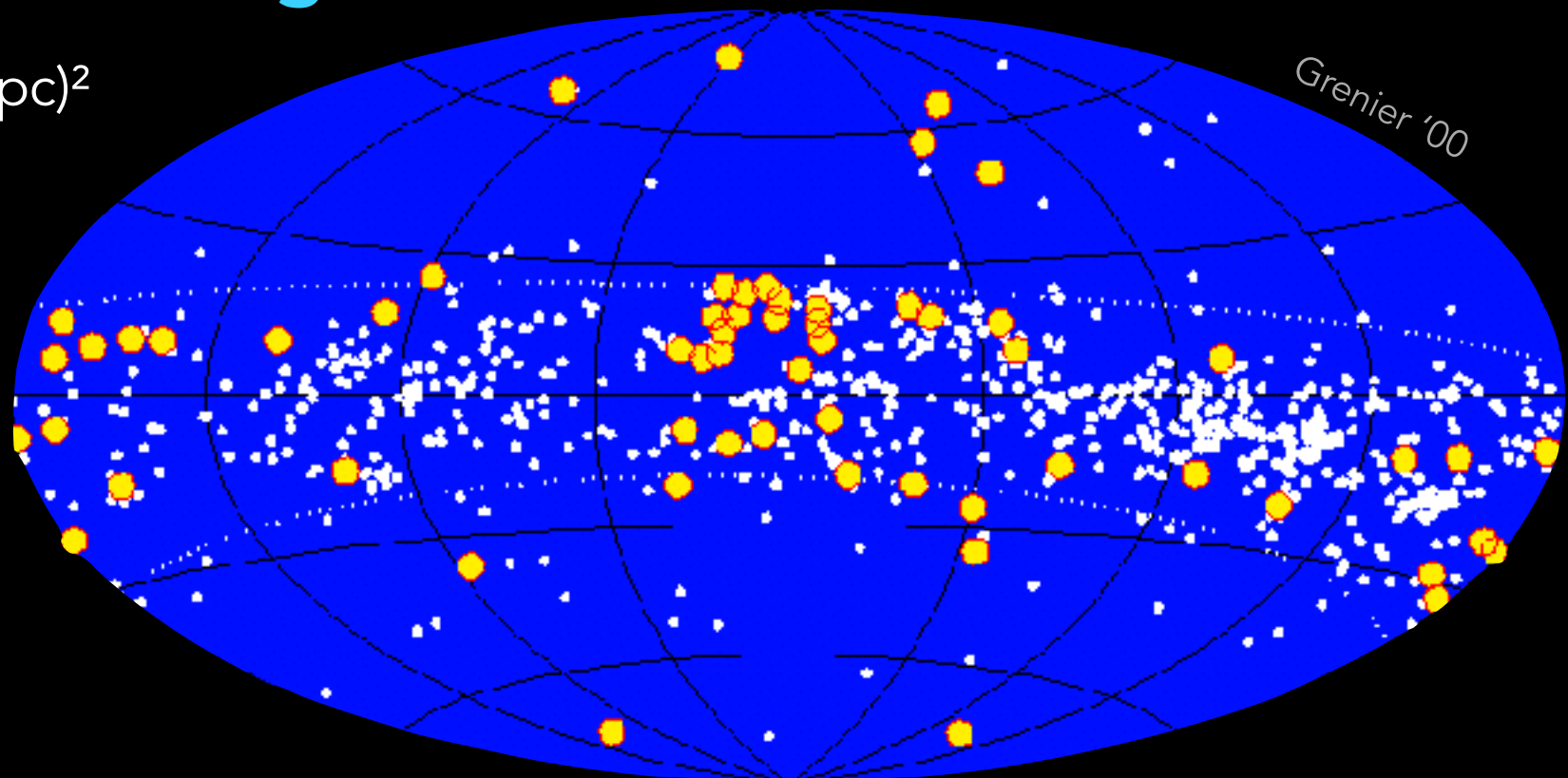
- off-beam pulsars ~ 20 over 5 Myr Harding & Zhang '01
but bright enough?

- BHoles accreting gas Dermer '01
but 3 to 9 times fewer than NS

- micro-quasars Kaufman Bernado+02
but too few

● hidden gas clumps ?

- but not resolved in HI & CO maps

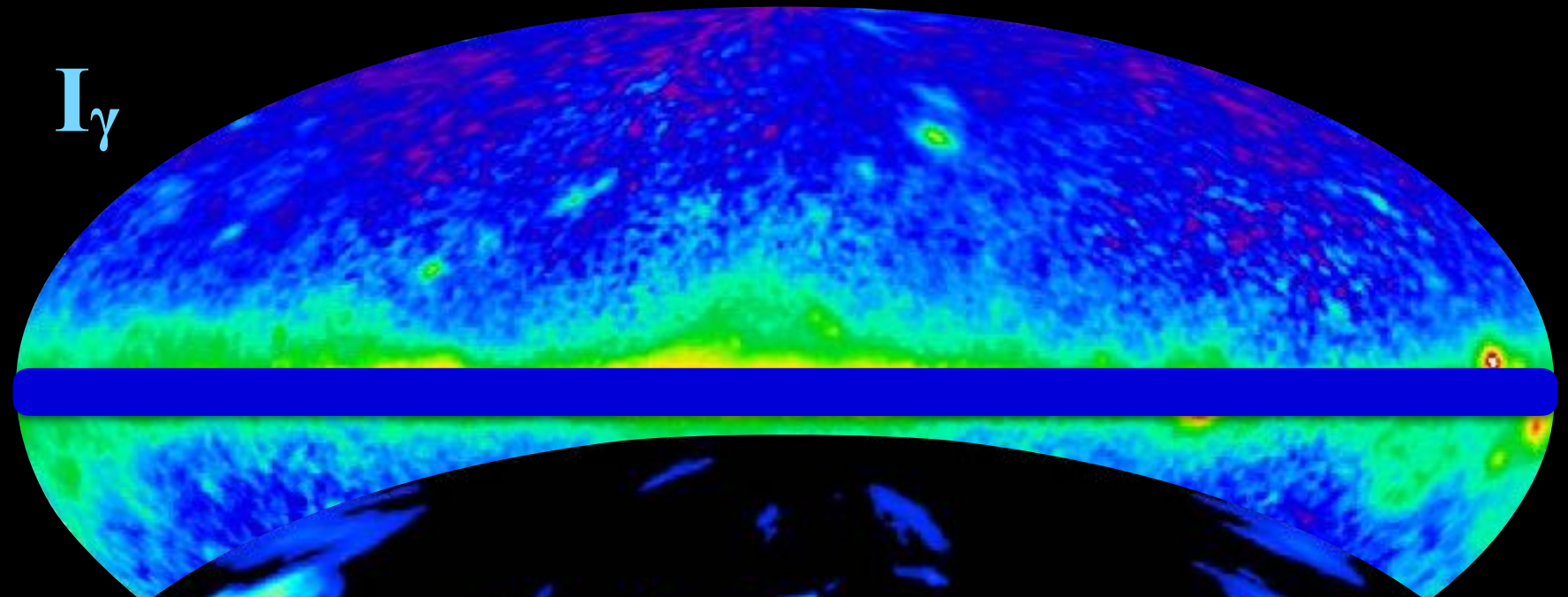


the dark neutral gas

● EGRET > 100 MeV



I_γ

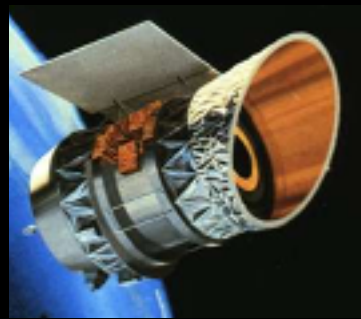


Hunter+ 97

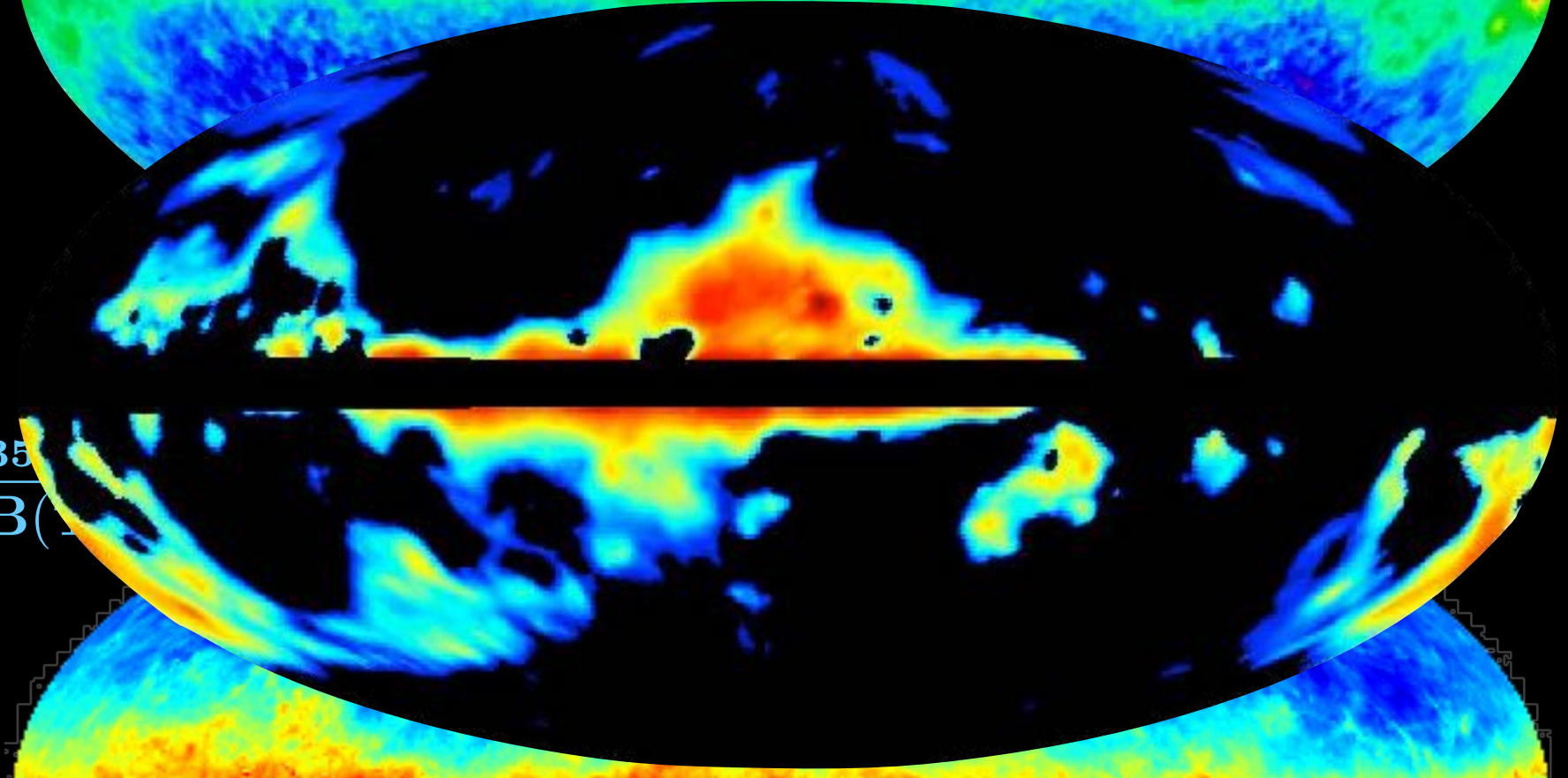
diffuse intensity \propto

$$\int n_{\text{gas}} n_{\text{CRs}} dl$$

● IRAS + DIRBE



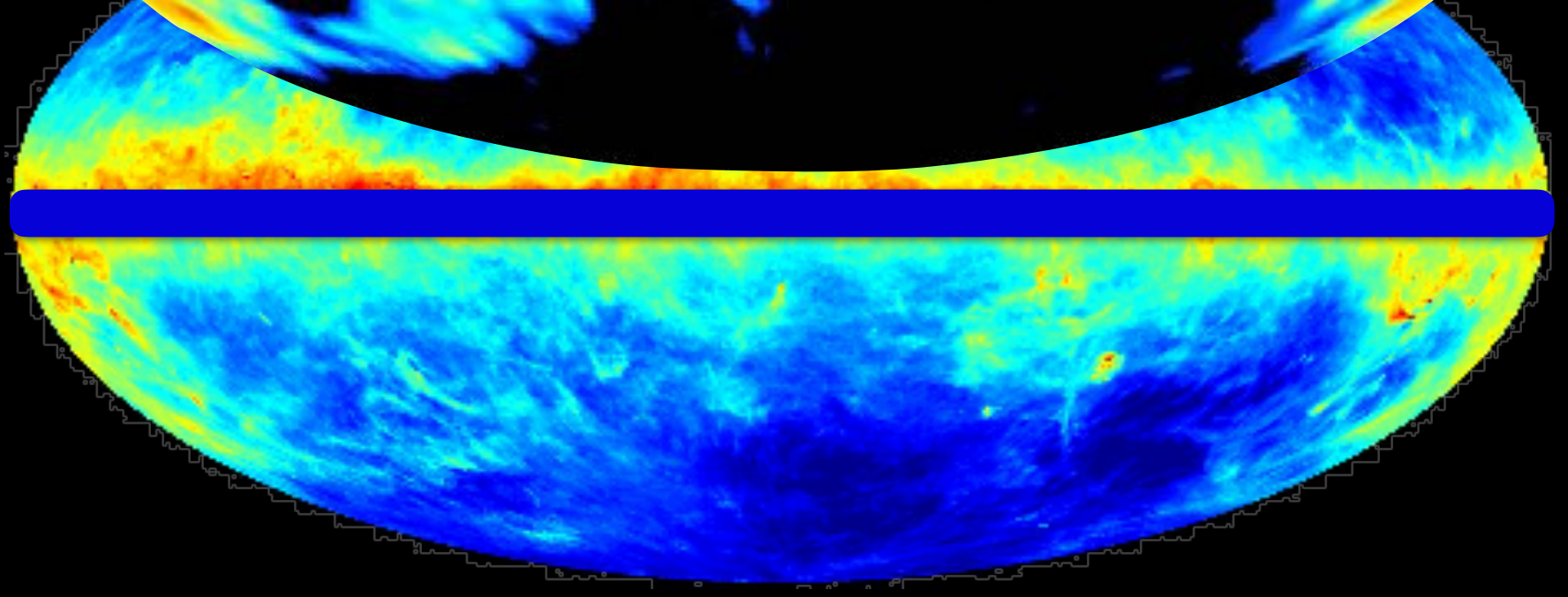
$$\tau_{353} = \frac{I_{353}}{B(\dots)}$$



Grenier+ 05

thermal intensity

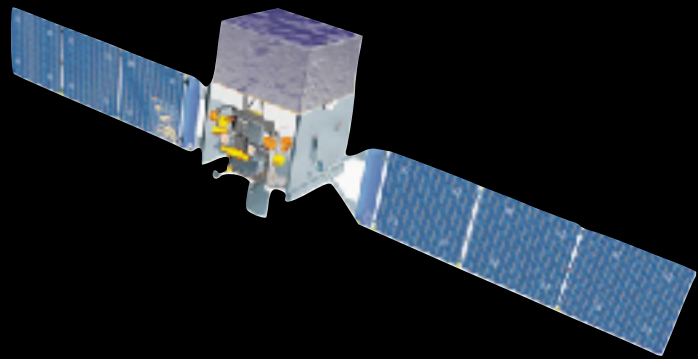
$$\int \frac{M_d}{M_g} n_g \kappa_0 \left(\frac{v}{v_0}\right)^\beta B_\nu(T_d) dl$$



Schlegel+ 98

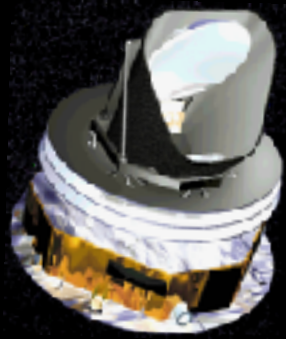
the dark neutral gas

Fermi-LAT > 0.6 GeV



diffuse intensity
 $\propto \int n_{\text{gas}} n_{\text{CRs}} dl$

Planck + IRAS

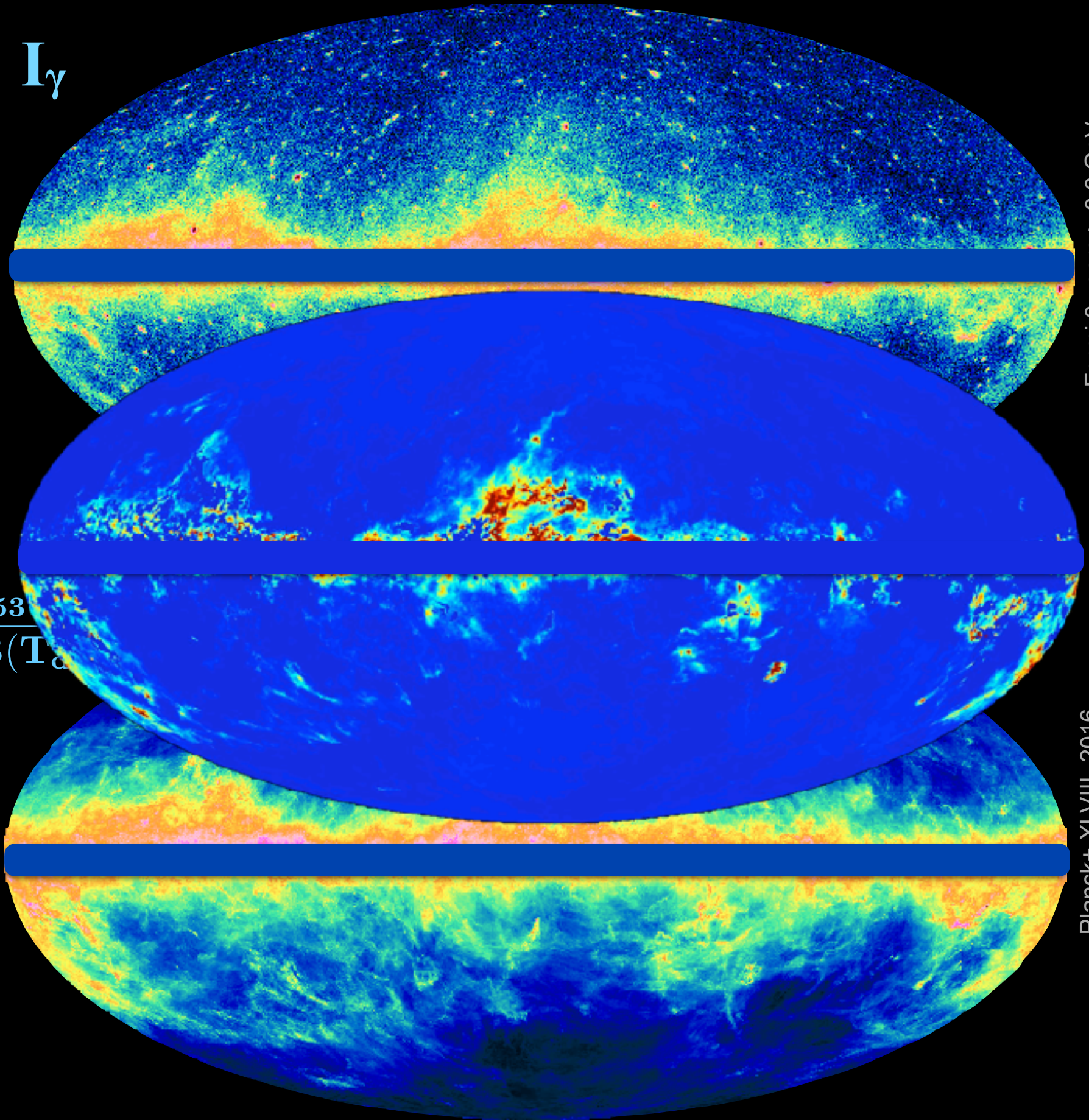


thermal intensity

$$\int \frac{M_d}{M_g} n_g \kappa_0 \left(\frac{v}{v_0} \right)^\beta B_\nu(T_d) dl$$

$$\tau_{353} = \frac{I_{353}}{B(T_d)}$$

I_γ



Fermi 8 years > 0.6 GeV

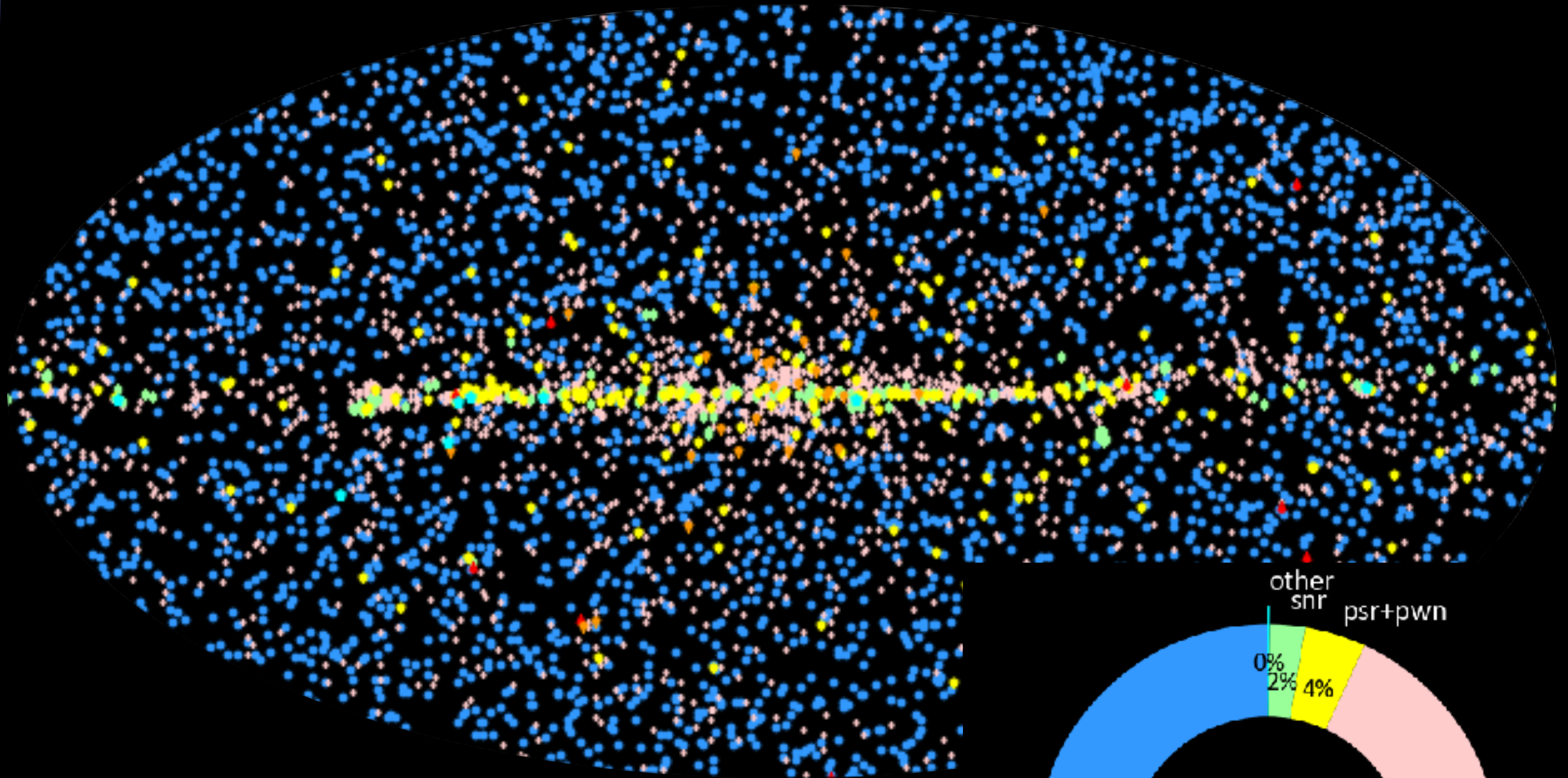
Planck+.XLVIII, 2016

setting GLAST on the right track

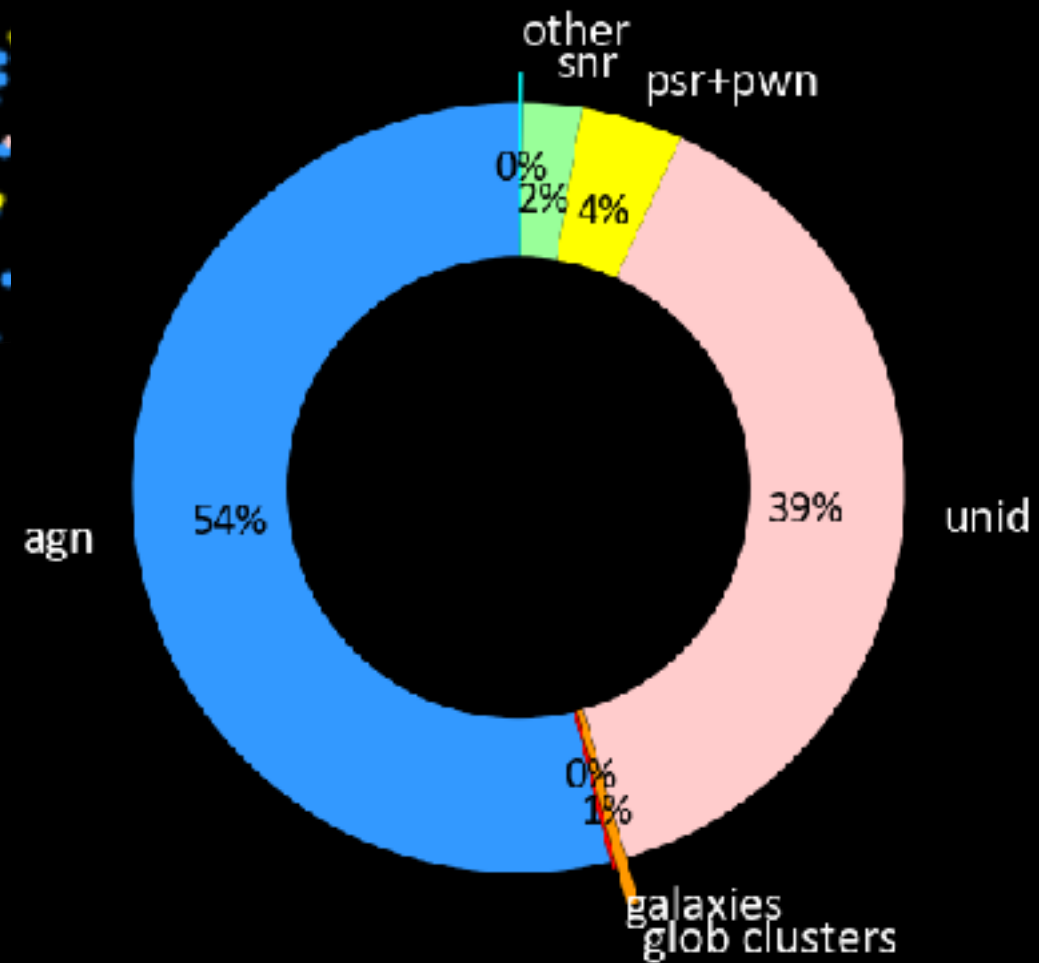
- chair of the Gamma Ray Astrophysics Program Working Group (GRAPWG) 1999
- led the GLAST project science requirement document (2000)
- helped with the science case for LAT, data right policy, deputy project scientist ...



more unidentified puzzles



- FL8Y source list = 5523 sources
 - localization $\approx 4.5'$



the efficiency of a broad smile, boldly looking forward



Neil 1952 - 2017

- graduated from University of Arizona
 - majored in music, then in physics
 - PhD Cal'Tech 1982 with Ed Stone
 - Voyager 1 + 2, Jupiter
 - data analysis from cosmic-ray detectors
 - O+S energetic ions (10-20 MeV/n) from Io accelerated at $r > 17 R_{\text{Jup}}$ in magnetosphere
- Ed would say that Neil was incredibly gifted

