

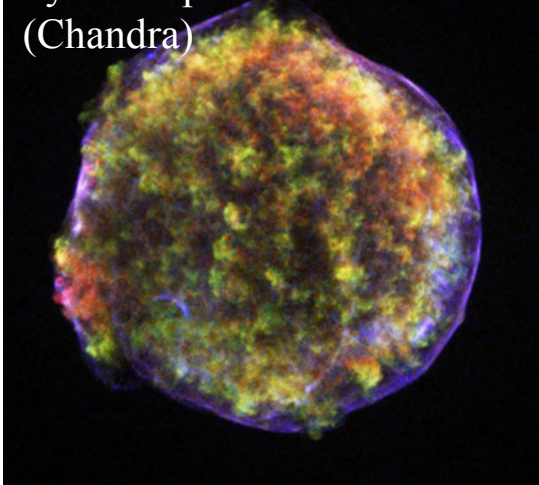
Constraining the Mass Loss Geometry of Beta Lyrae

Jamie Lomax

Collaborators: Jennifer Hoffman,
Nick Elias and Bruce Holenstein

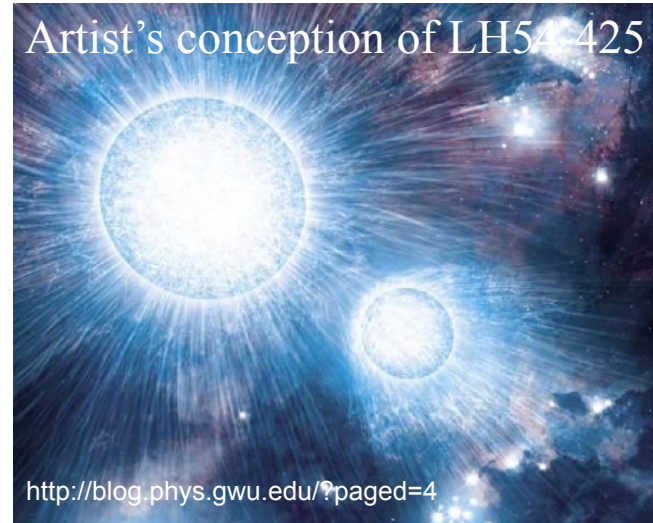
Mass Loss in Binary Systems

Tycho Supernova Remnant
(Chandra)



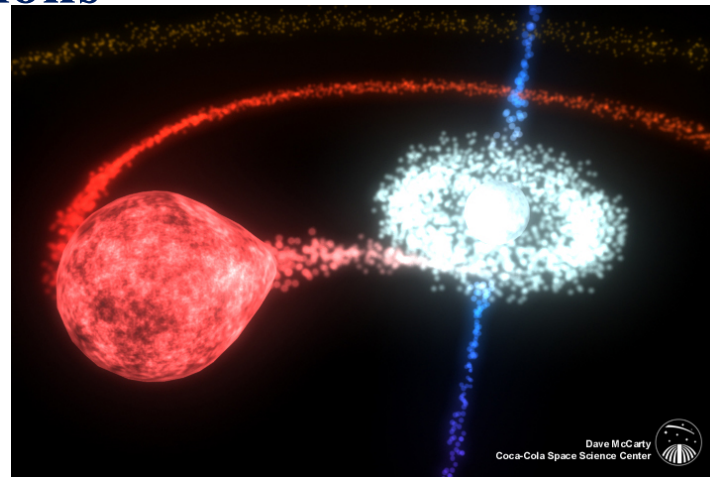
Explosions

Artist's conception of LH54-425



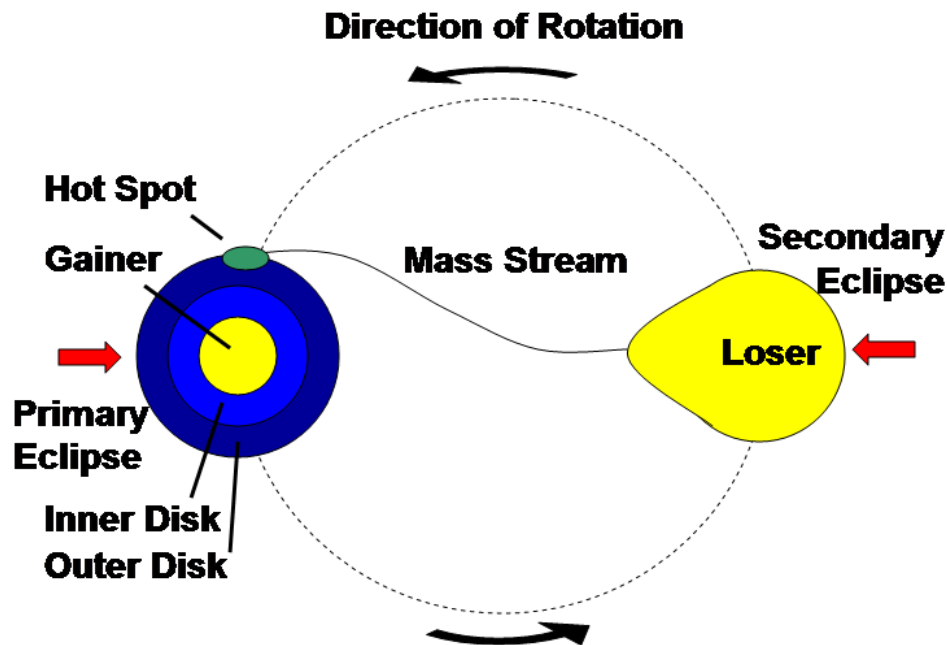
<http://blog.phys.gwu.edu/?paged=4>

Stellar Winds

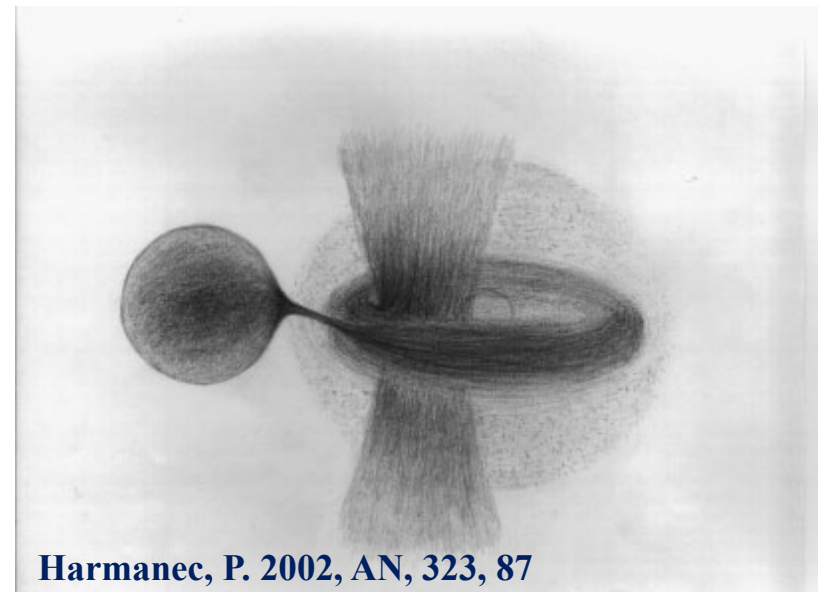


Roche lobe transfer

Beta Lyrae: What We Know



Polar View

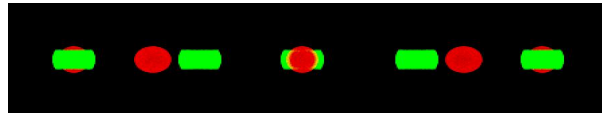
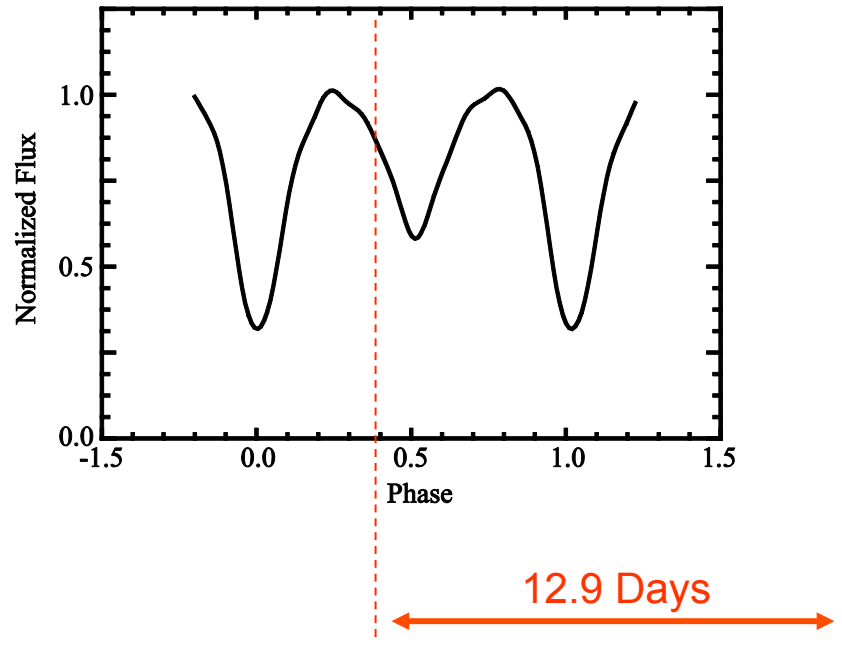


Harmanec, P. 2002, AN, 323, 87

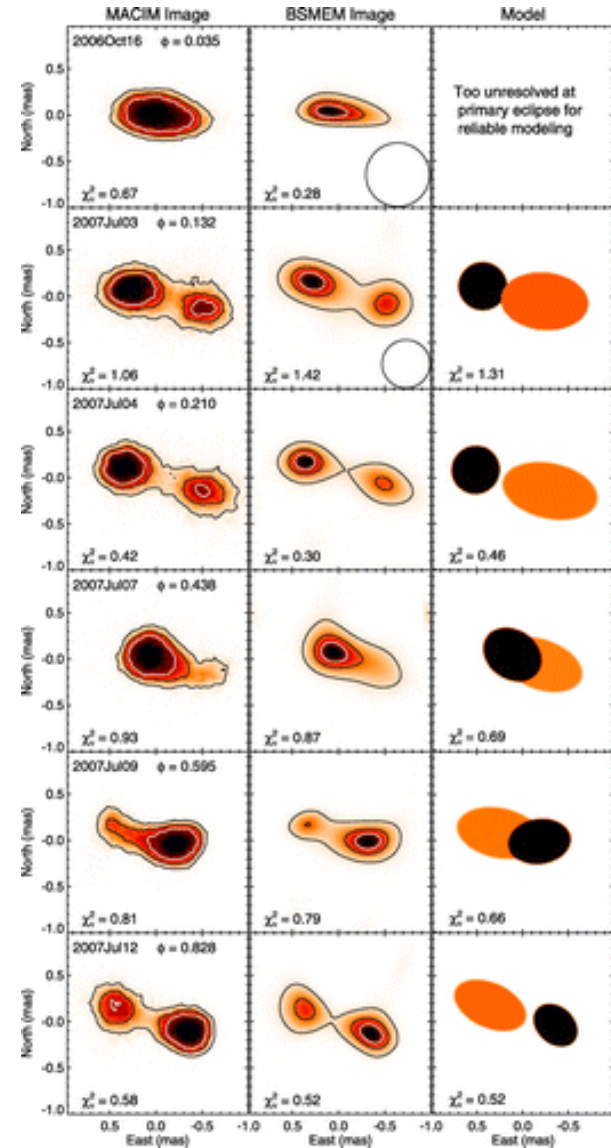
Edge On View

Beta Lyrae: What We Know

V Band Light Curve



IR H band: 1.5 to 1.8 μ m



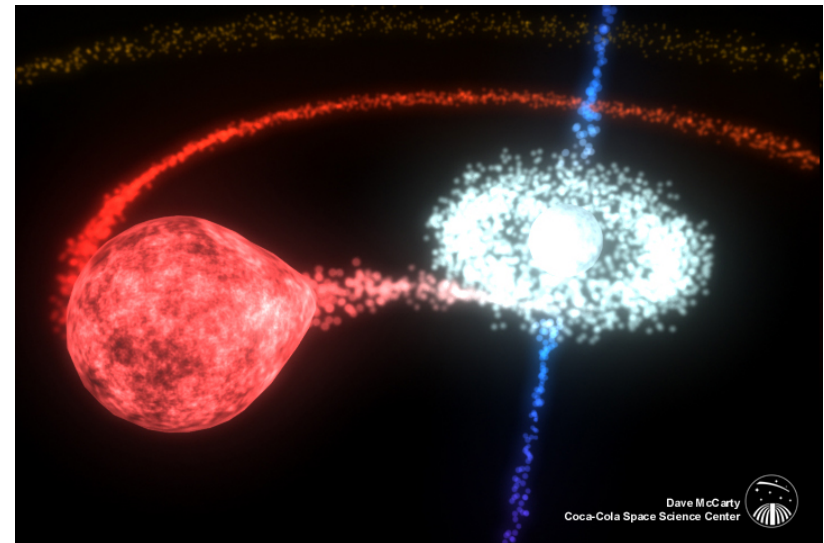
Harmanec et al., 1996, A&A, 312, 879
 Zhao et al., 2008, ApJ, 684, L95

Beta Lyrae: What We Don't Know

Where is the mass lost?

How much matter is in the jets?

What are the characteristics of the jets and mass stream?

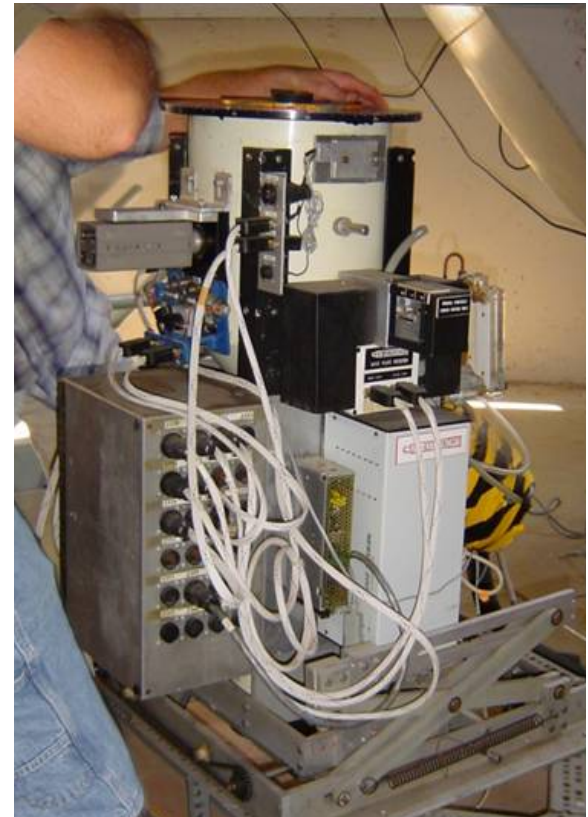


The Data Set

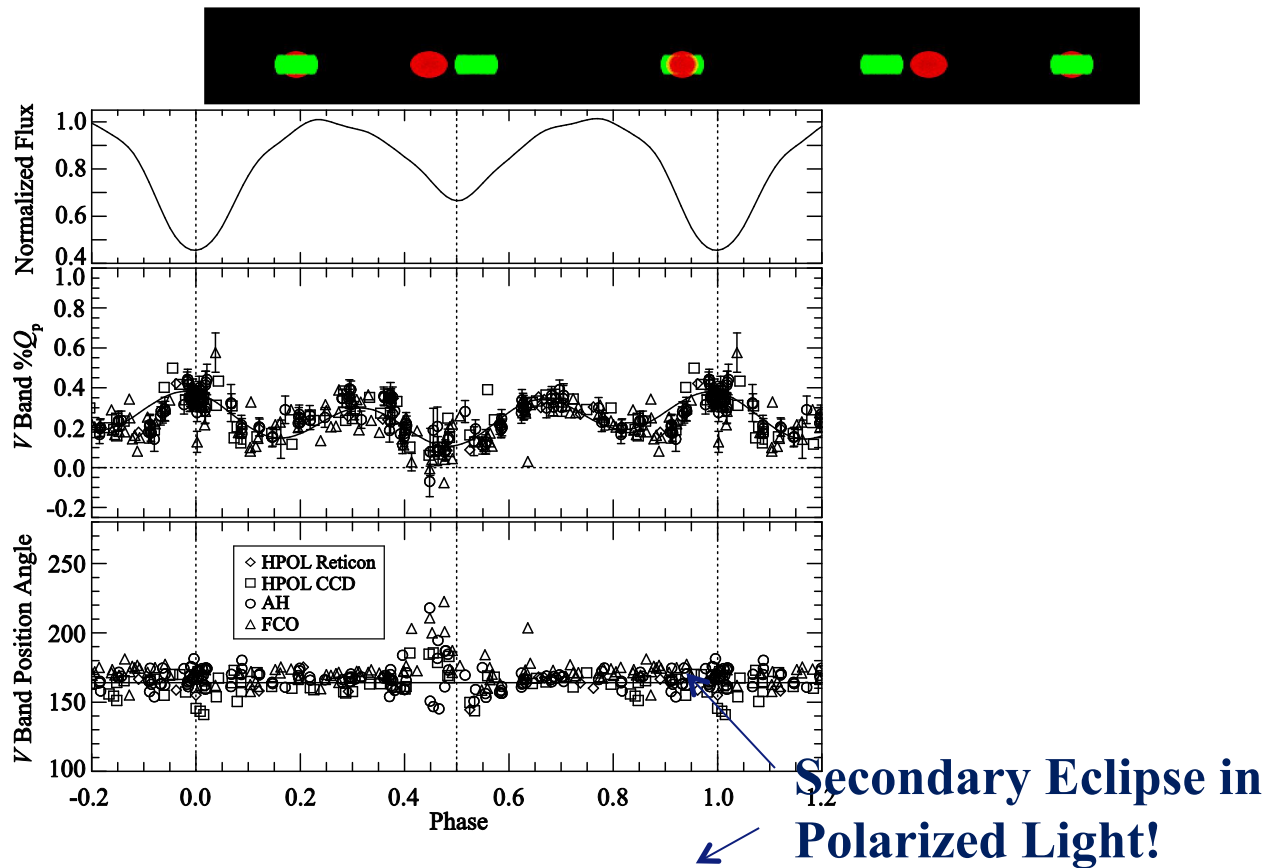
**HPOL: 14 Reticon (3200-7500Å)
53-55 CCD (3200-10,500Å)**

**Appenzeller & Hiltner (1969): 37 *B*
127 *V***

**Flower & Cook Observatory: 19 *B*
88 *V*
17 *R***

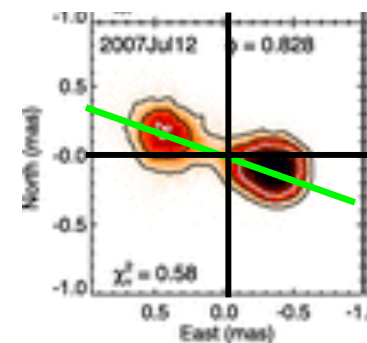


V Band Polarized Light Curve

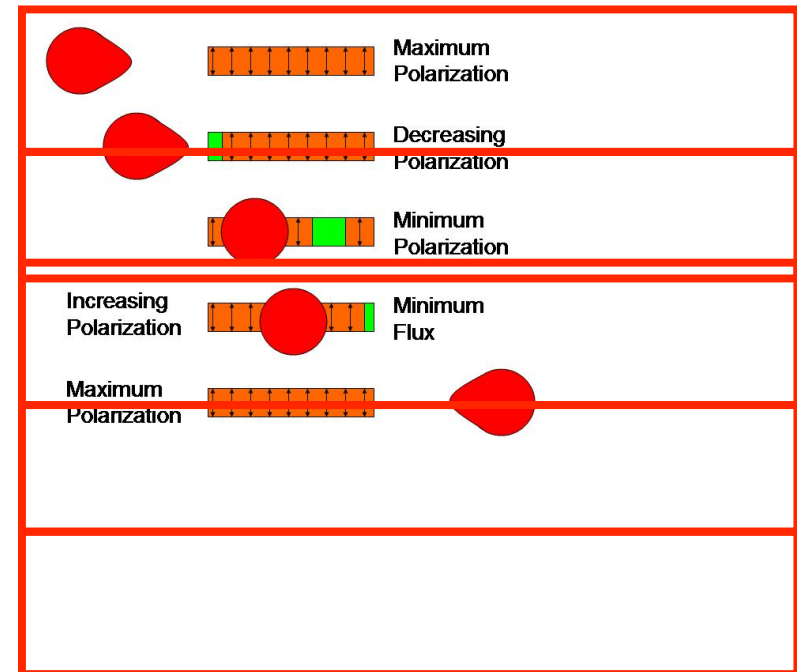
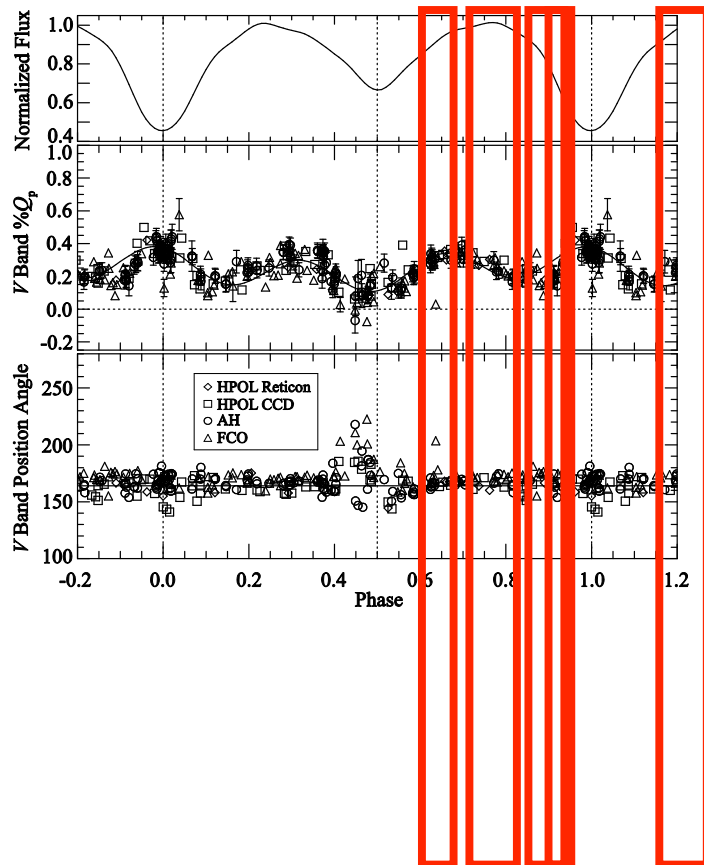


Zhao et. al. (2008):
 $253^\circ \pm 1.97^\circ$
 $251^\circ \pm 1.83^\circ$

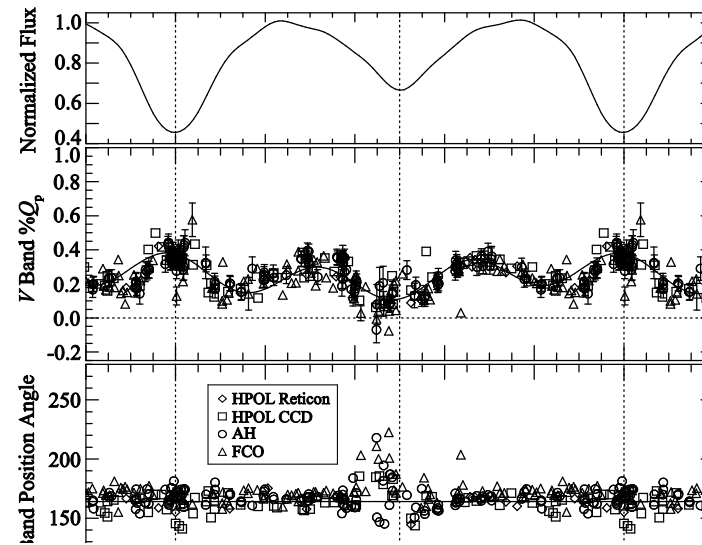
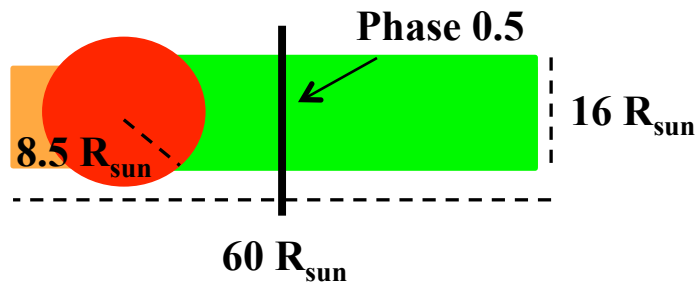
HPOL Observations:
 $164^\circ (\pm 1^\circ) + 90^\circ = 254^\circ$



Proposed Geometry

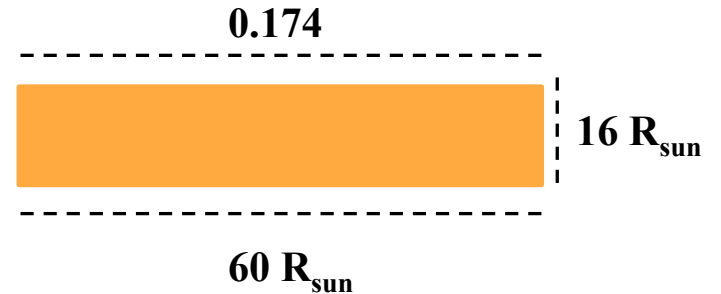
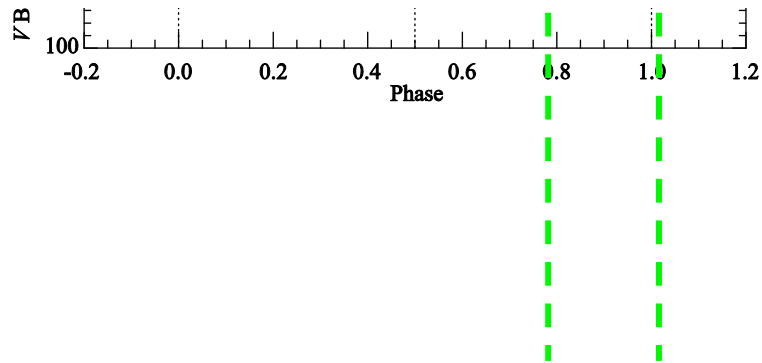


Hot Spot $\%Q_p$ Size Estimate



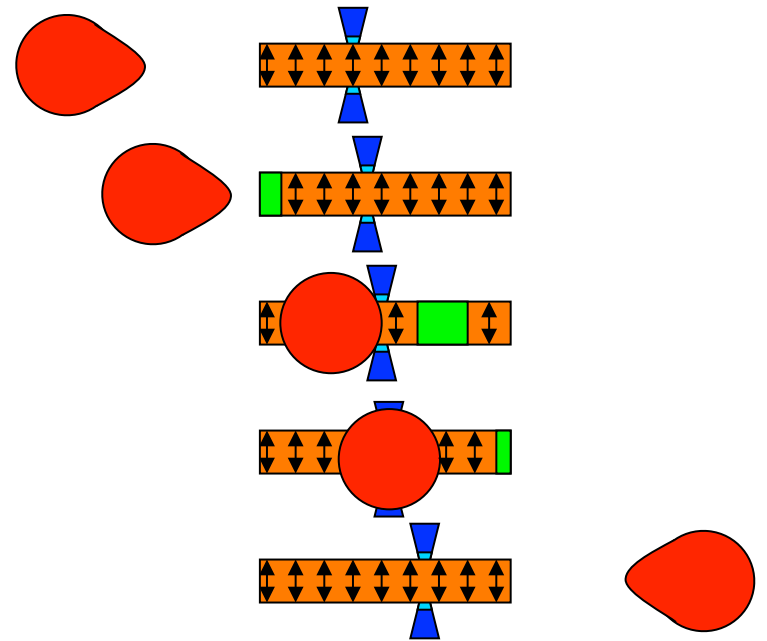
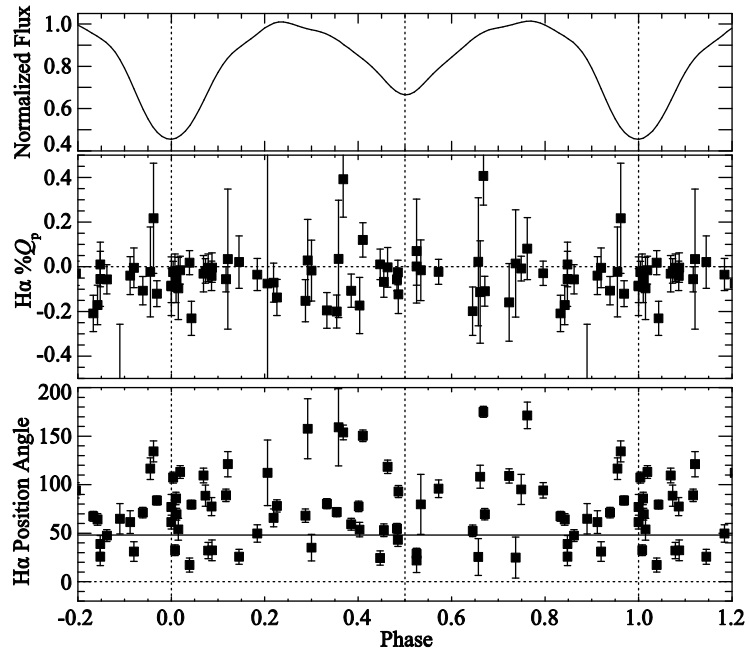
Band	2 nd ary Phase	$\%Q_p$ Hot Spot Size
B	0.449	$28 R_{\text{sun}}$
V	0.457	$28 R_{\text{sun}}$
R	0.482	$40 R_{\text{sun}}$
I	0.455	$40 R_{\text{sun}}$

Hot Spot PA Size Estimate



Band	Size in Phase	Max Hot Spot Size
B	0.14	48 R_{sun}
V	0.16	55 R_{sun}
R	0.125	43 R_{sun}
I	0.09	31 R_{sun}

H α Line



Conclusions

Band	PA Hot Spot Size	%Q _p Hot Spot Size
B	48 R _{sun}	28 R _{sun}
V	55 R _{sun}	28 R _{sun}
R	43 R _{sun}	40 R _{sun}
I	21 R _{sun}	40 R _{sun}

Line polarimetry suggests origin of jets is not hot spot

Do other Roche lobe-transfer systems show secondary eclipse effect?

What does the jet location mean in terms of evolution?

Future Work:

X-ray data

Line polarization

Better monitoring near secondary eclipse

Period Analysis

Acknowledgements:

Jennifer Hoffman

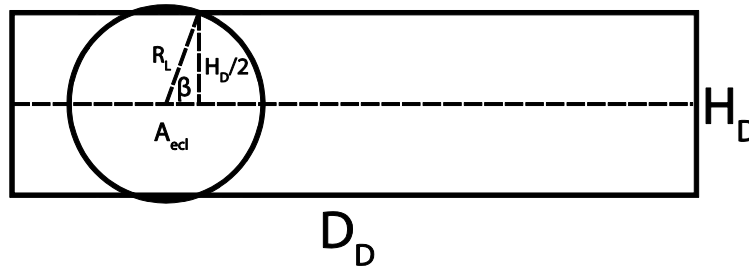
Marilyn R. Meade

Ken Nordsieck

Brian Babler

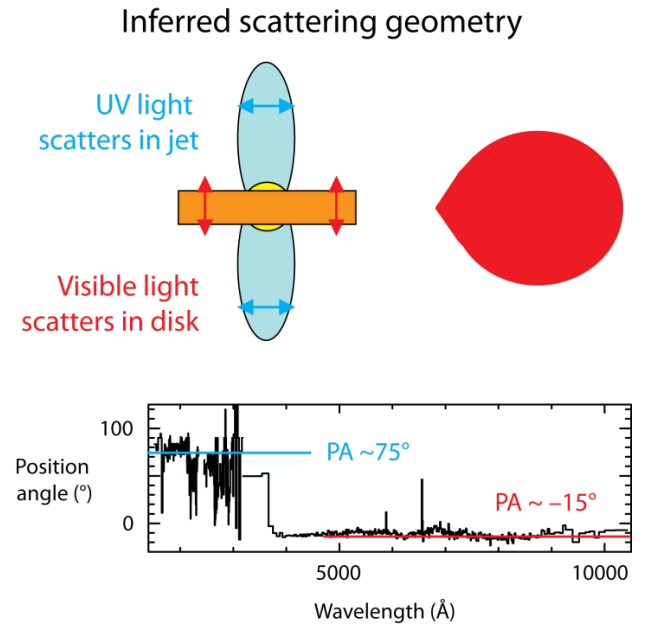
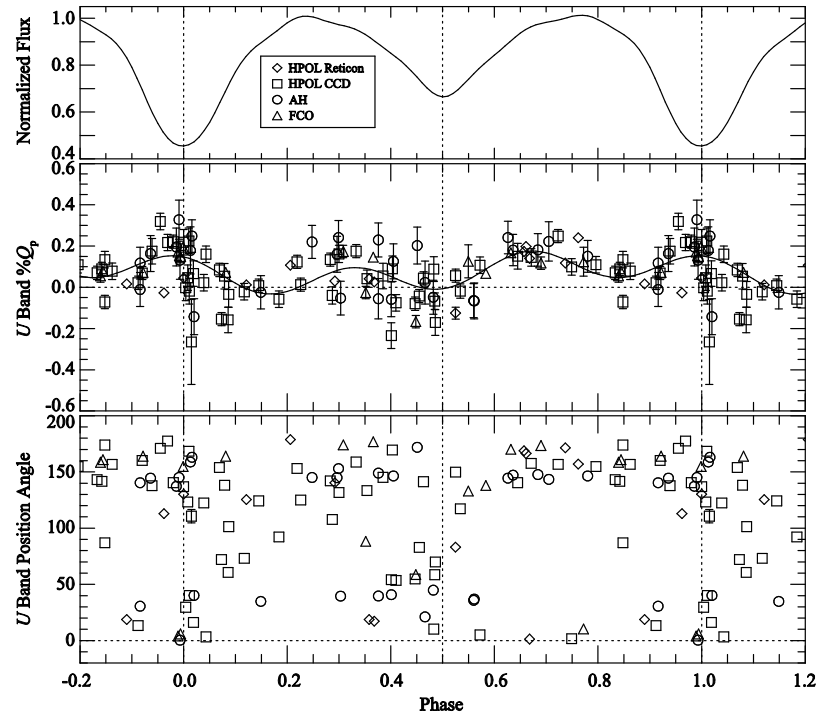
Nick Elias

Hot Spot Simple Model Size Estimate

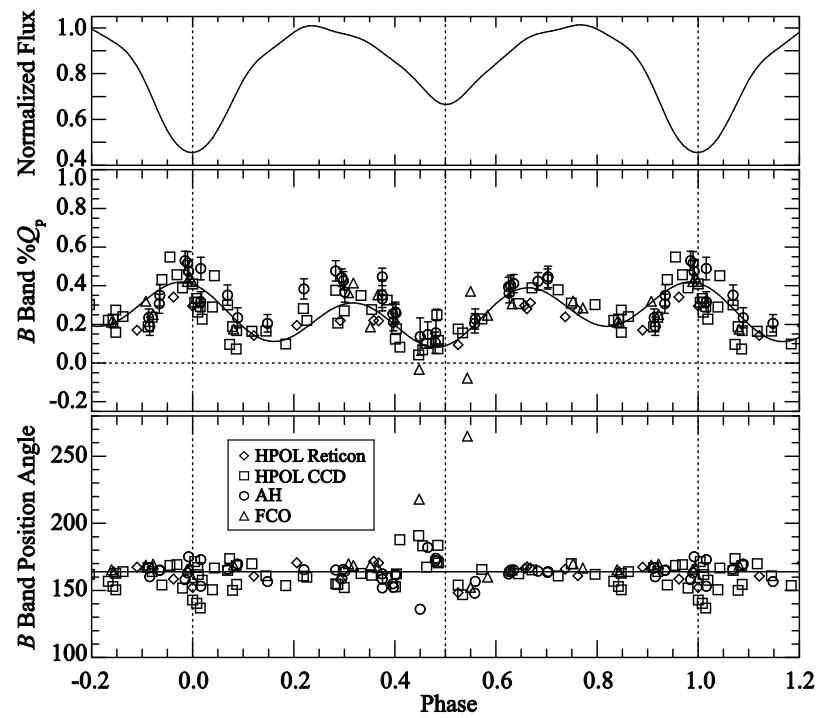


Band	q_{DC}	q_{min}	HS_{SM}
B	0.187	0.041	37
V	0.179	0.081	23
R	0.171	0.074	24
I	0.140	0.076	19

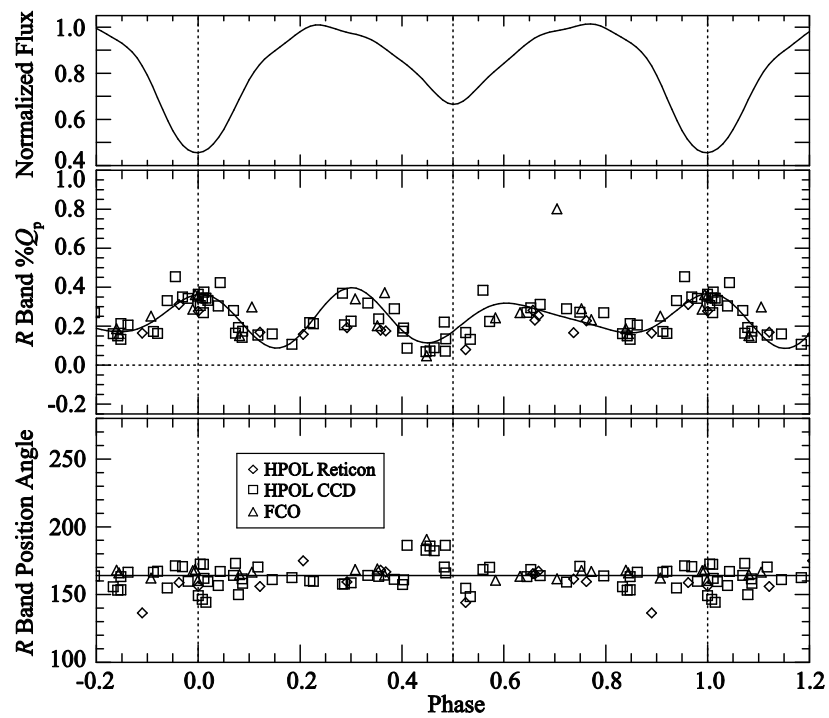
U Band



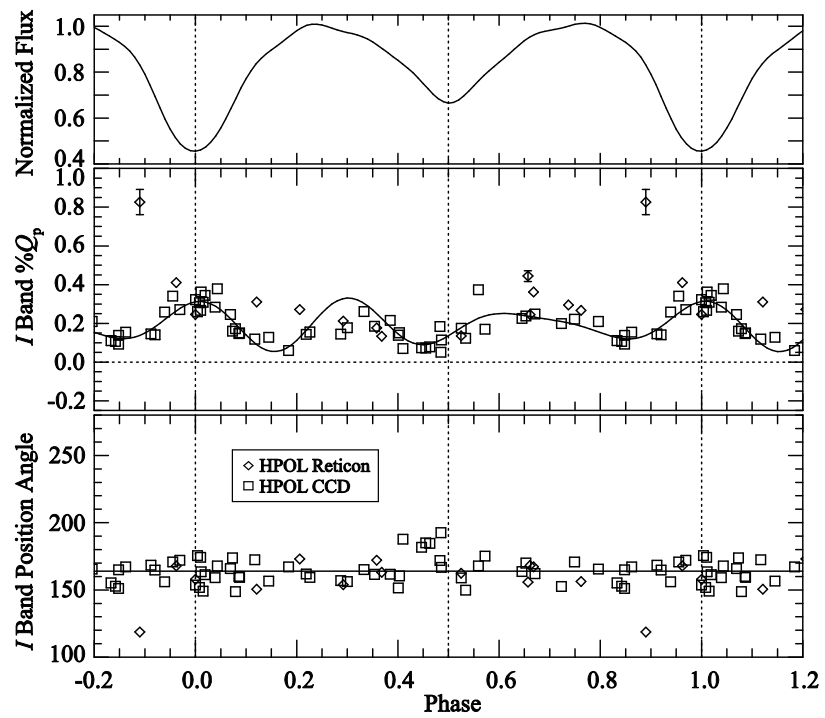
B Band



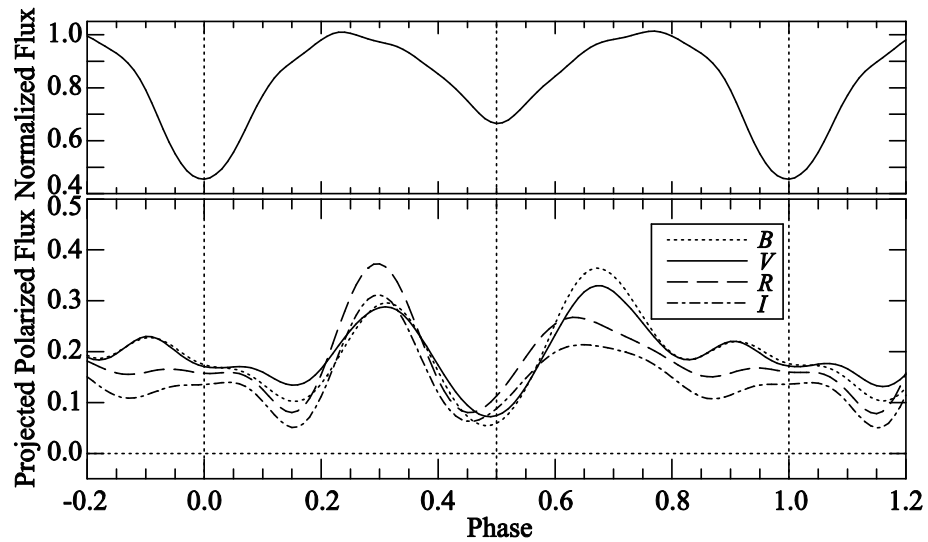
R Band



I Band



Projected Polarized Flux



V Band Polarization and % U_p

